

## Summary

- 1. Laser amplifiers
  - 1. Pulse amplification
  - 2. ASE, Parasitic Lasing, Thermal effects
  - 3. Amplifier chains, Regenerative amplifiers, multipass amplifiers
  - 4. Nonlinear effects: B-integral
- Ultrashort amplifiers
  - 1. Short pulse propagation
  - 2. CPA
  - 3. Ti:Sapphire amplifiers

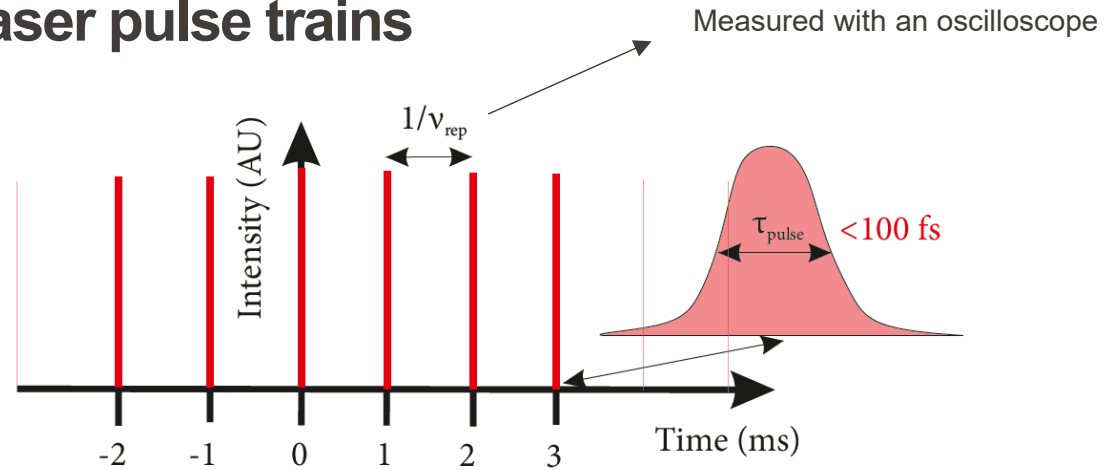
# Advanced Radiation Sources - PHSY761

## Lecture 05

8 October  
2024

# Laser pulse trains

Train of equally-spaced pulses:  
The «repetition rate» is the  
inverse of the temporal pulse  
separation («duty cycle»)



Pulse duration: we define it as the FWHM of the intensity

Oscilloscope (ns)  
Special methods are needed for  
shorter pulse

Average power: energy averaged over a time  $T \gg 1/v_{\text{rep}}$

Measure of the energy carried in  
time: measured with a slow power  
meter

Pulse power: pulse energy / pulse duration

Measure of the **peak** power of  
the pulse (even very moderate  
average powers can have huge  
pulse power)

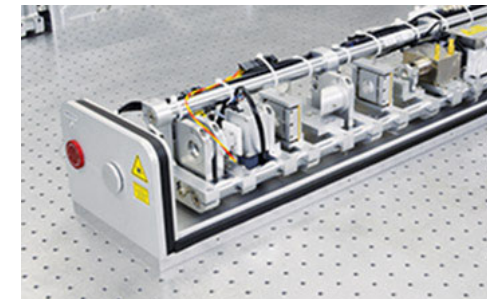
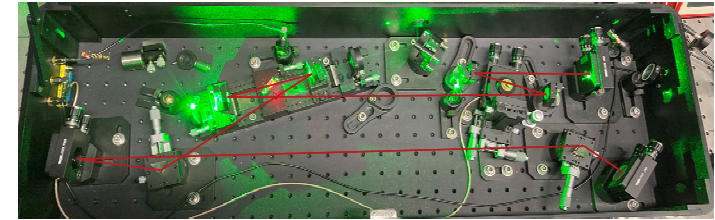
Pulsed laser sources we have seen so far:

Typical specifications of laser oscillators  $E(v_{\text{rep}})$ :

Mode-locked fs Ti:Sapphire lasers ( $\approx 10$  fs)  $\rightarrow \approx 10$  nJ (100 MHz)

Q-switched Nd laser ( $\approx 10$  ns)  $\rightarrow \approx 100$  mJ (Hz - kHz)

These correspond to MW pulses  $\rightarrow$  Different strategies are required for higher peak intensities



MW  $\rightarrow$  GW  $\rightarrow$  TW  $\rightarrow$  PW

Perturbative  
Nonlinear optics

Strong field effects

Relativistic  
nonlinear optics

PEAK INTENSITY

$10^{13} \text{ W/cm}^2$

$10^{18} \text{ W/cm}^2$

$\hookrightarrow$  LINKED TO  
MODE QUALITY

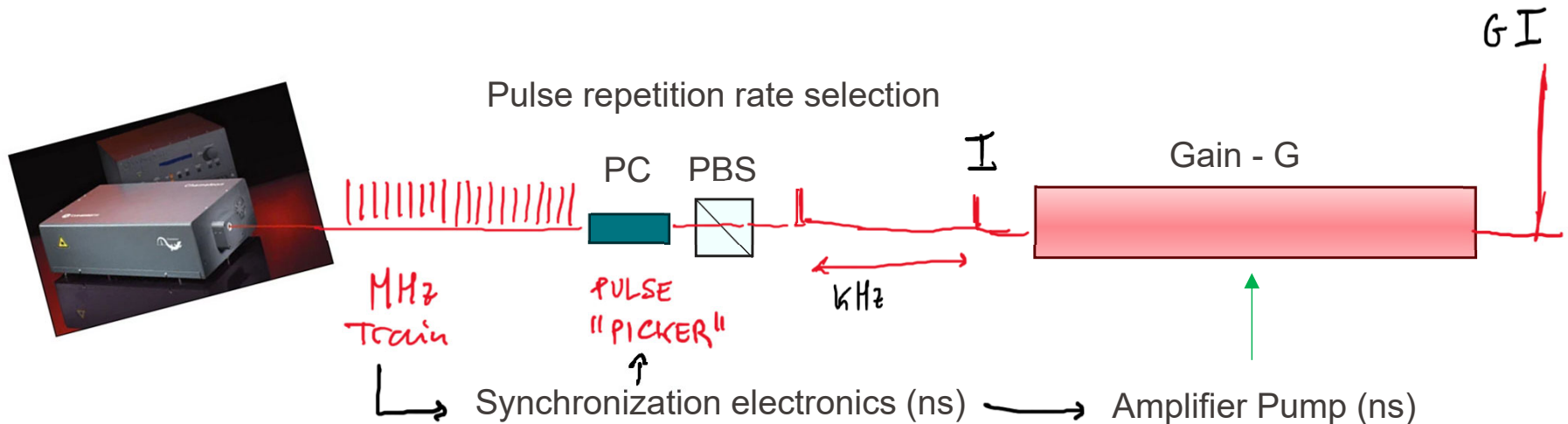
$\rightarrow$  DIFFRACTION LIMITED SOURCES

-

Why so hard to increase the power of a femtosecond oscillator?

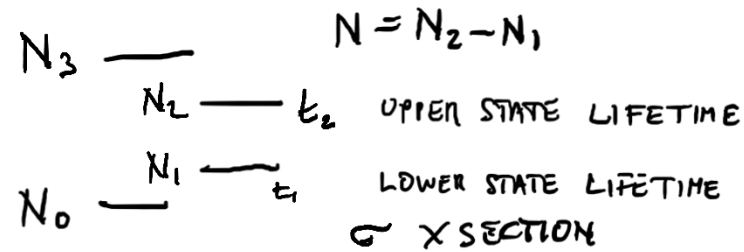
- Lack of suitable pump lasers
- High intracavity power and intensity: thermal and nonlinear effects.

## Master Oscillator / Power Amplifier (MOPA)

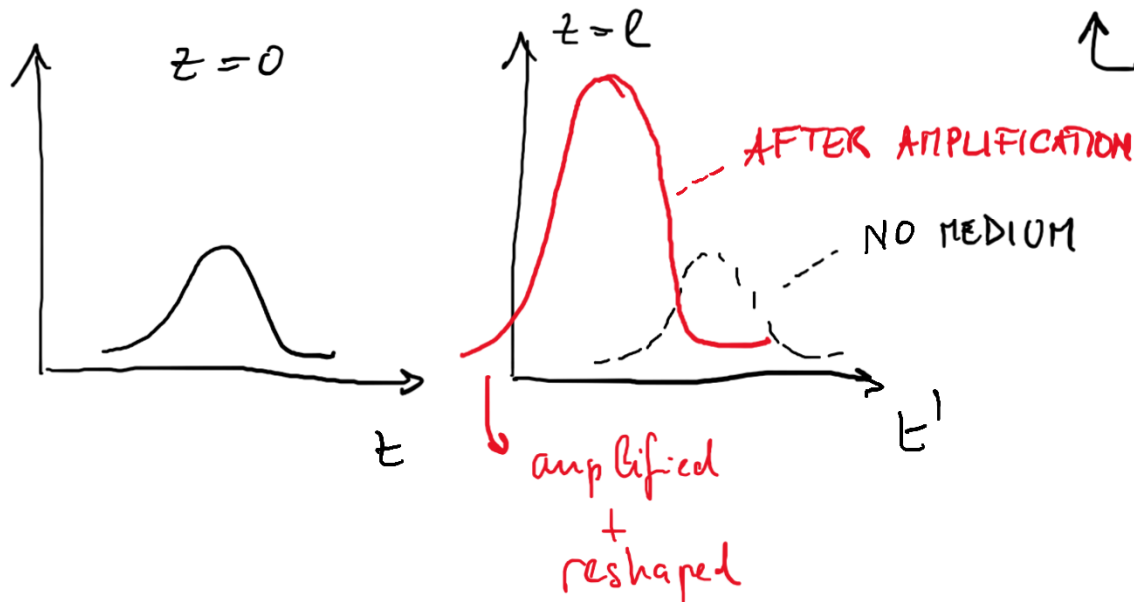
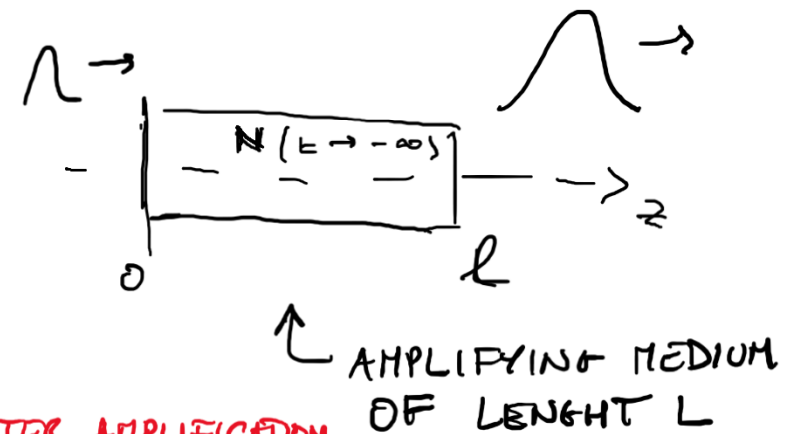


- 4-LEVEL SYSTEM
- $\tau_p$  pulse duration

$$t_1 \ll \tau_p \ll t_2$$



1d MODEL  $\rightarrow$  WE WANT TO KNOW  
 $\rightarrow I(z, t)$  ;  $N(z, t) \leftarrow$   
 INTENSITY POP. INVERSION



RATE EQ.  $\rightarrow$  FOR POP. INVERSION

$$\frac{\partial N(z,t)}{\partial t} = \overset{\substack{\text{Transition} \\ \text{rate}}}{W} N(z,t) + \dots$$

NEGLLECT OTHER TRANSITIONS MODIFYING INVERSION

$$= \phi \sigma N = \frac{I}{h\nu} \sigma N$$

$$= \frac{I(z,t) N(z,t)}{I_s}$$

$I_s = \frac{h\nu}{\sigma}$   
SATURATION EN. DENSITY  
[J/cm<sup>2</sup>]

$$I_s \approx 1 \text{ J/cm}^2$$

Ti: Sapphire

$$I_s = 1 - 2 \text{ mJ/cm}^2$$

Dye laser

E.M. energy density  
↓

$$\frac{\partial \mathcal{E}}{\partial t} = \left. \frac{\partial \mathcal{E}}{\partial t} \right|_1 + \left. \frac{\partial \mathcal{E}}{\partial t} \right|_2 + \left. \frac{\partial \mathcal{E}}{\partial t} \right|_3$$

STIMULATED  
PROCESSES

LOSSES  
(e.g. defects...)

photon  
flux in/out  
 $dV = S dz$

$$\left( \frac{\partial \mathcal{E}}{\partial t} \right)_1 = \omega N \hbar \omega$$

$$= \sigma I N$$

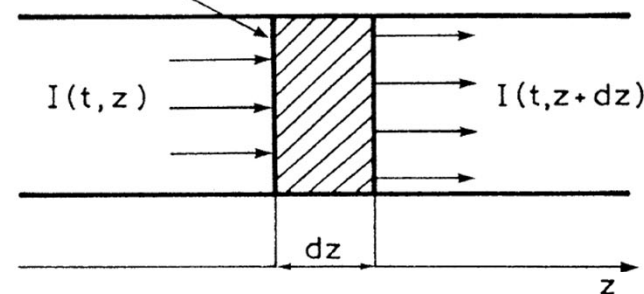
$$\left( \frac{\partial \mathcal{E}}{\partial t} \right)_3 S dz = S (I(t, z+dz) - I(t, z))$$

$$\left( \frac{\partial \mathcal{E}}{\partial t} \right)_2 = -\alpha_L I$$

↳ due to ampl. coeff...

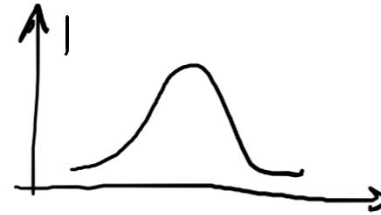
$$\frac{\partial \mathcal{E}}{\partial t} = \frac{1}{c} \frac{\partial I}{\partial t} = \sigma I N - \alpha_L I + \frac{\partial I}{\partial z}$$

Cross-sectional  
area  $S$



$$\begin{aligned} \textcircled{1} & \left\{ \begin{aligned} \frac{1}{c} \frac{\partial I}{\partial t} + \frac{\partial I}{\partial z} &= \sigma N I - \alpha I \\ \textcircled{2} & \left\{ \begin{aligned} \frac{\partial N}{\partial t} &= -\frac{N I}{P_s} \end{aligned} \right. \end{aligned} \right. \end{aligned}$$

CAN BE NUMERICALLY INTEGRATED -  
EXACT SOLUTION DEPENDS ON THE INPUT  
SHAPE



$$I(t \rightarrow \pm\infty) = 0$$

PULSE  
ENERGY  
density

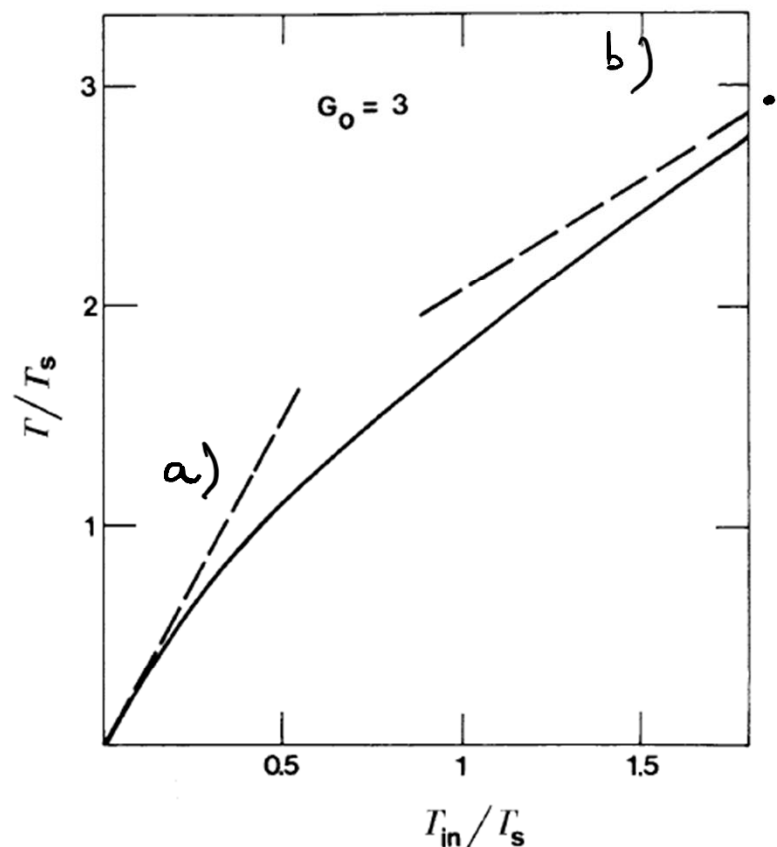
$$P = \int_{-\infty}^{+\infty} I dt$$

SOME GENERAL CONSIDERATION CAN BE  
DRAWN FOR THE ENERGY FLUENCE  $P$

$$\textcircled{2} \rightarrow N(t \rightarrow +\infty) = N(t \rightarrow -\infty) e^{-P/P_s} = N_0 e^{-P/P_s} \quad \left( \text{INTEGRATE } \int_{-\infty}^{+\infty} dt \right)$$

$$\textcircled{1} \int_{-\infty}^{+\infty} dt \quad \text{BOTH SIDES} \quad ; \quad \text{INSERT } \textcircled{2}; \quad I(\infty) = 0$$

$$\frac{dP}{dz} = g P_s \left[ 1 - \exp\left(-P/P_s\right) \right] - \alpha P \quad g = N_0 \sigma \quad \text{UNSATURATED GAIN COEFFICIENT}$$



ASSUMING  $\alpha_L \approx 0$  HAS ANALYTIC SOL.

$$\Gamma(l) = \Gamma_s \ln \left[ 1 + G_0 \left[ \exp\left(\frac{\Gamma_{in}}{\Gamma_s}\right) - 1 \right] \right]$$

$$G_0 = e^{-g l} \quad \text{UNSATURATED GAIN (SMALL-SIGNAL GAIN)} \quad \Gamma_{in} = \Gamma(0)$$

Two extreme cases.   
 Linear...  $\ln(1+x) \sim x$

$$a) \frac{\Gamma_{in}}{\Gamma_s} \ll 1 \Rightarrow \Gamma(l) = G_0 \Gamma_{in}$$

LINEAR REGIME

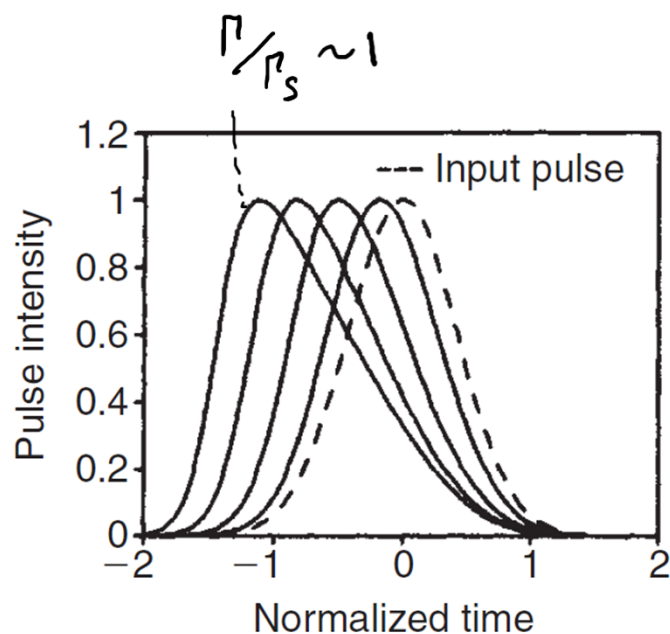
$$b) \frac{\Gamma_{in}}{\Gamma_s} \gg 1 \Rightarrow \Gamma(l) = \Gamma_{in} + g l \Gamma_s$$

DEEP SATURATION REGIME

MAXIMUM ENERGY EXTRACTION

$$g \Gamma_s l = \underbrace{N_0 h \nu l}_{\text{ALL THE ATOMS CONTRIBUTE}}$$

## Transverse mode and temporal reshaping during amplification

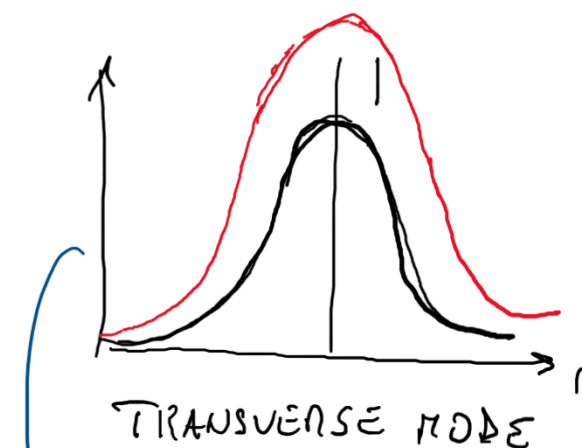


NUMERICAL  
RESULTS FOR :

GAUSSIAN INPUT

$$10^{-4} < P/P_s < 1$$

$$G_0 = 10^4$$



SATURATION AMPLIFY

THE WINGS MORE THAN  
THE PEAK

IN GENERAL  $\rightarrow$  SPATIAL PROFILE  
WORSEN

What limits the maximum fluence which can be extracted by an amplifier?

- Medium losses  $\longrightarrow \Gamma_{max} \rightarrow \frac{g\Gamma_s}{\alpha_L} \quad (\alpha \ll g)$
- Optical damage  $\longrightarrow \Gamma_d \approx 10 \text{ J/cm}^2 \quad \Gamma \approx g\ell\Gamma_s < \Gamma_d$

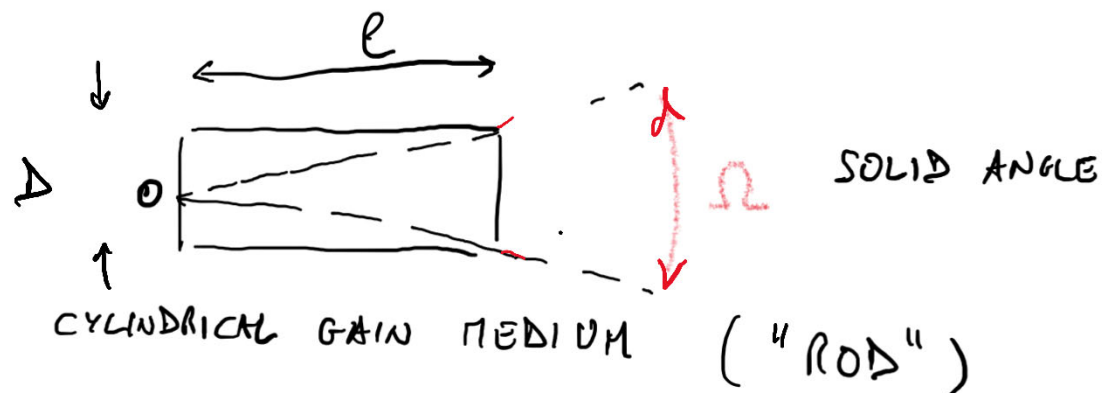
Effects arising due to a high gain  $G_0$

- Amplified spontaneous emission (ASE)
- Parasitic lasing

Limitation due to high average power and high intensity

- Thermal effects  $\longrightarrow$  Mode distortion
- Nonlinear effects  $\longrightarrow$  Temporal distortion

- Amplified spontaneous emission (ASE)



$$G = \exp[\sigma(N_2 - N_1)l] \quad \text{gain (assume to be large e.g. } 10^4 \text{)}$$

$\Rightarrow$  STRONG AMPLIFICATION OF THE FLUORESCENCE  
FROM  $\odot$

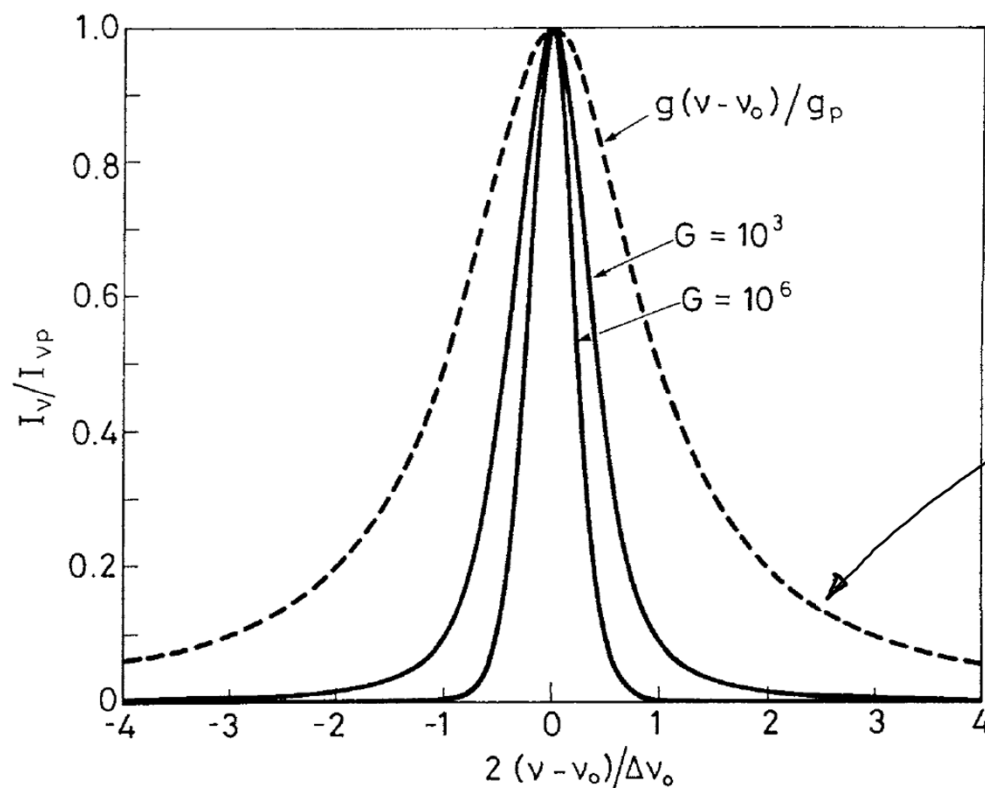
$$D \ll l \quad \Omega \sim \frac{\pi D^2}{4l^2}$$

"MIRRORLESS LASER"

SUPERLUMINESCENCE

- NARROWER THAN SPONT. EMISSION (SPECTRALLY) MORE "COHERENT"  
↳ gain narrowing
- DIRECTIONAL ( $\Omega$ )
- SOFT "THRESHOLD"

## SPECTRAL NARROWING OF THE PULSE

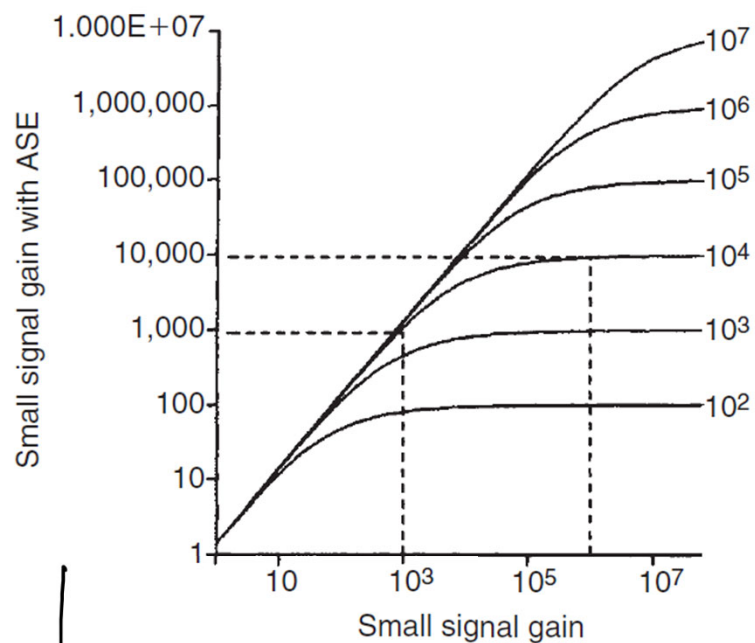
TRANSITION LINE  
SHAPE

$$G_0 = \exp(g \ell)$$

$$\underbrace{\sigma(\omega) N_0 \ell}$$

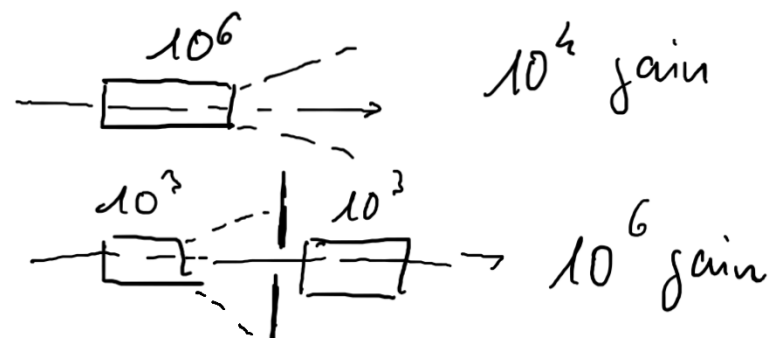
$$G(\omega) = \exp[\sigma(\omega) N_0 \ell]$$

## Mitigation of ASE:



$$G_0 = e^{\sigma N_s L}$$

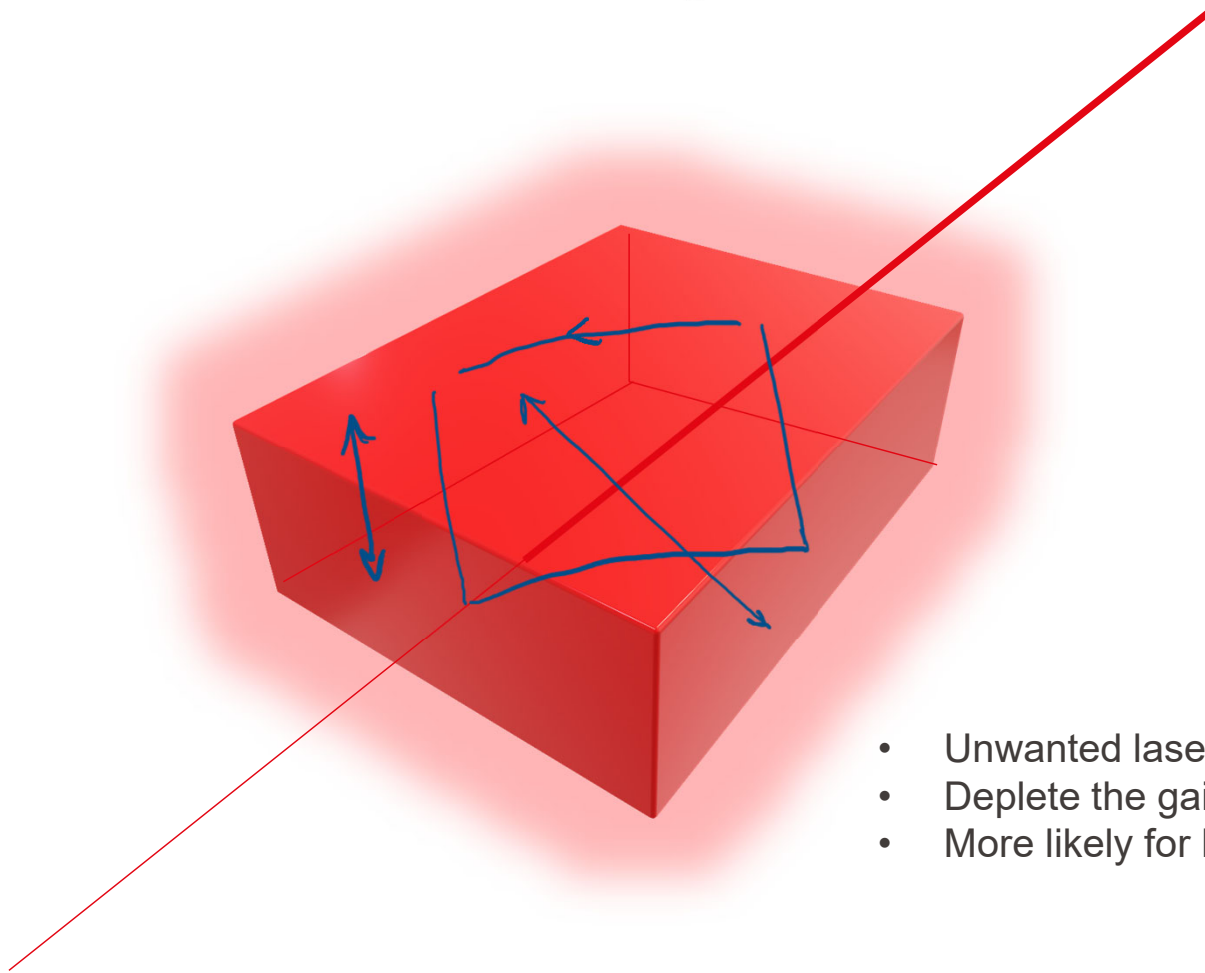
IDEA: 2X GAIN SECTIONS



Actual  
gain considering  
ASE

Theoretical  
gain w/o ASE

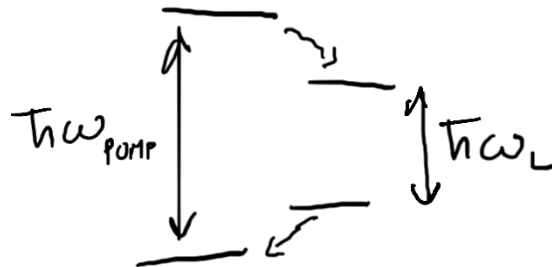
## Parasitic lasing



- Unwanted laser oscillations
- Deplete the gain and limit the maximum
- More likely for high G

## Thermal effects

Quantum defect in a lasing transition



$$h(\omega_p - \omega_L) = \delta > 0$$

HAS TO BE DISSIPATED  
AS HEAT

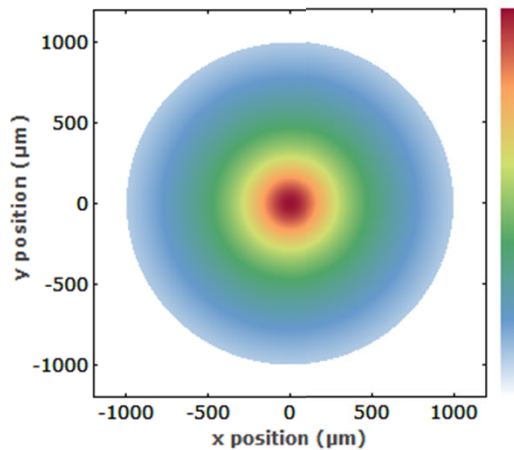
NON RADIATIVE  
SCATTERING WITH THE  
LATTICE

Thermo-optical effects  
Thermo-elastic effects

$$\frac{\partial n}{\partial T}$$

Thermal "lensing"

STRAIN  
NON UNIFORM  
PUMP!

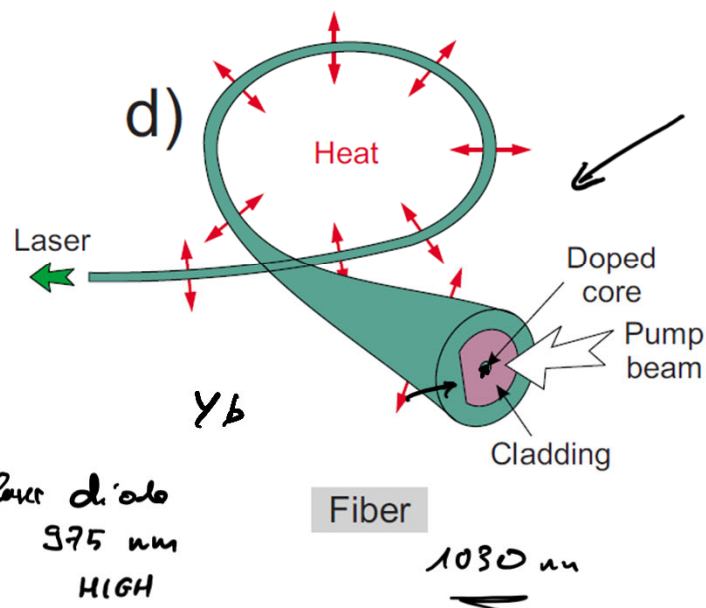
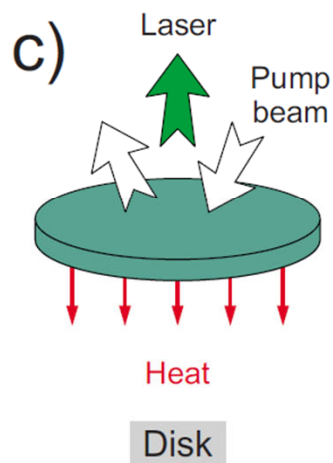
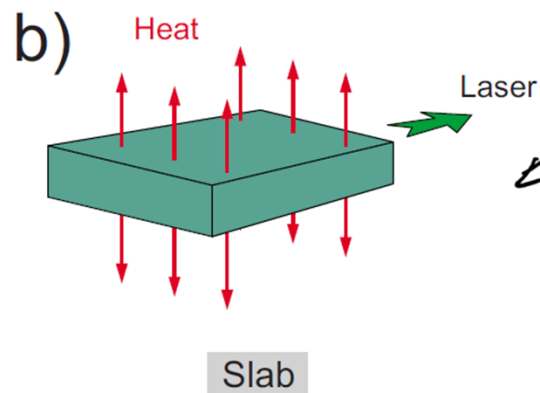
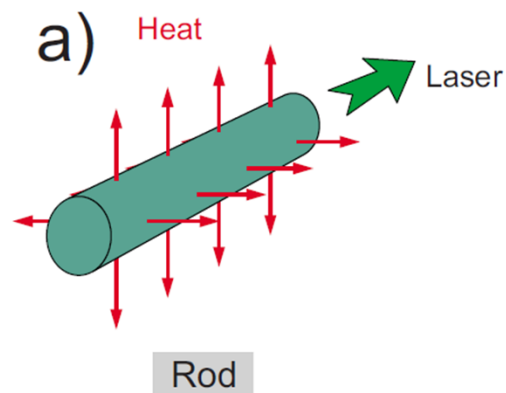


T DISTRIBUTION ON END-PUMPED ROD

$$f_{TH}^{-1} = \frac{\partial n}{\partial T} \frac{1}{2kA} P_{heat}$$

CAN BE ACCOUNTED FOR IN THE DESIGN  
 Thermal conductivity  
 dissipated power

## Shape of the amplifying medium:



• BETTER COOLING

• 1d Flow cylindrical Thermal lens

MOSTLY IMMUNE TO THERMAL EFFECTS

• kW CW LASERS

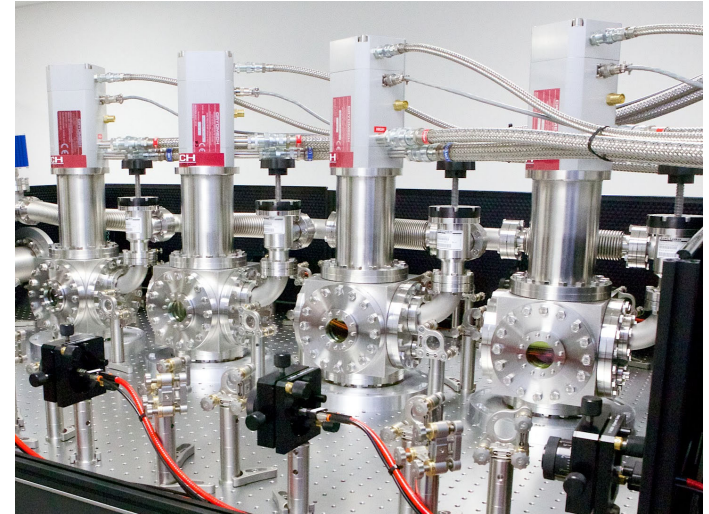
• SMALL CORE (nonlinear effects large)

• 1d flow  
• very efficient cooling  
- low gain

TABLE I  
SAPPHIRE PROPERTIES AT 300 AND 77 K<sup>a</sup>

| Property                                               | At 300 K                                | At 77 K                                 |
|--------------------------------------------------------|-----------------------------------------|-----------------------------------------|
| Thermal Conductivity, $\kappa$                         | $0.33 \text{ W cm}^{-1} \text{ K}^{-1}$ | $10 \text{ W cm}^{-1} \text{ K}^{-1}$   |
| Dependence of Refractive Index on Temperature, $dn/dT$ | $1.28 \times 10^{-5} \text{ K}^{-1}$    | $0.19 \times 10^{-5} \text{ K}^{-1}$    |
| Thermal Expansion Coefficient, $\alpha$                | $5.0 \times 10^{-6} \text{ K}^{-1}$     | $0.34 \times 10^{-6} \text{ K}^{-1}$    |
| Volumetric Heat Capacity, $c\rho$                      | $3.1 \text{ J cm}^{-3} \text{ K}^{-1}$  | $0.25 \text{ J cm}^{-3} \text{ K}^{-1}$ |

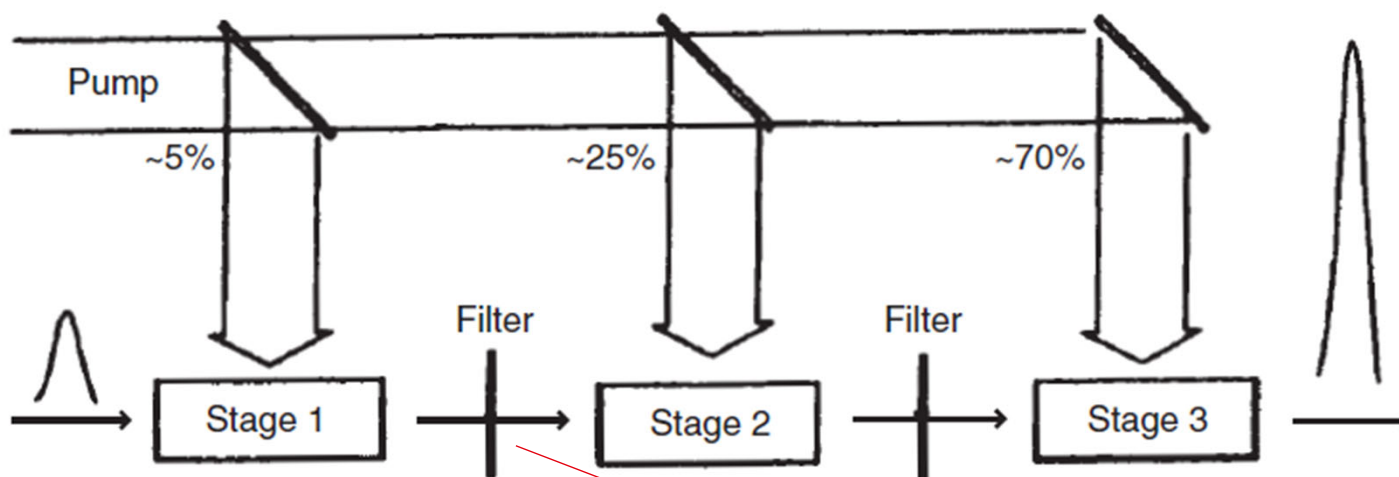
IEEE JOURNAL OF QUANTUM ELECTRONICS, VOL. 21, NO. 4, APRIL 1991



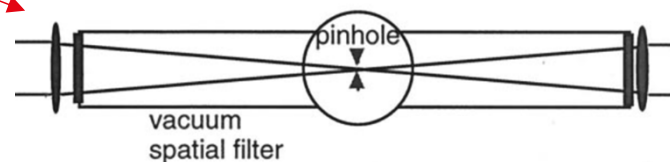
Thor-300 Cryogenic Yb:YAG Laser Amplifier System

- Improved thermal parameters
- Closed loop cryostats

**Amplifier chains:** Given a target total amplification, better to separate the amplification in several stages



- High G pre-amplifiers
- Low G power amplifiers
- Spatial filters



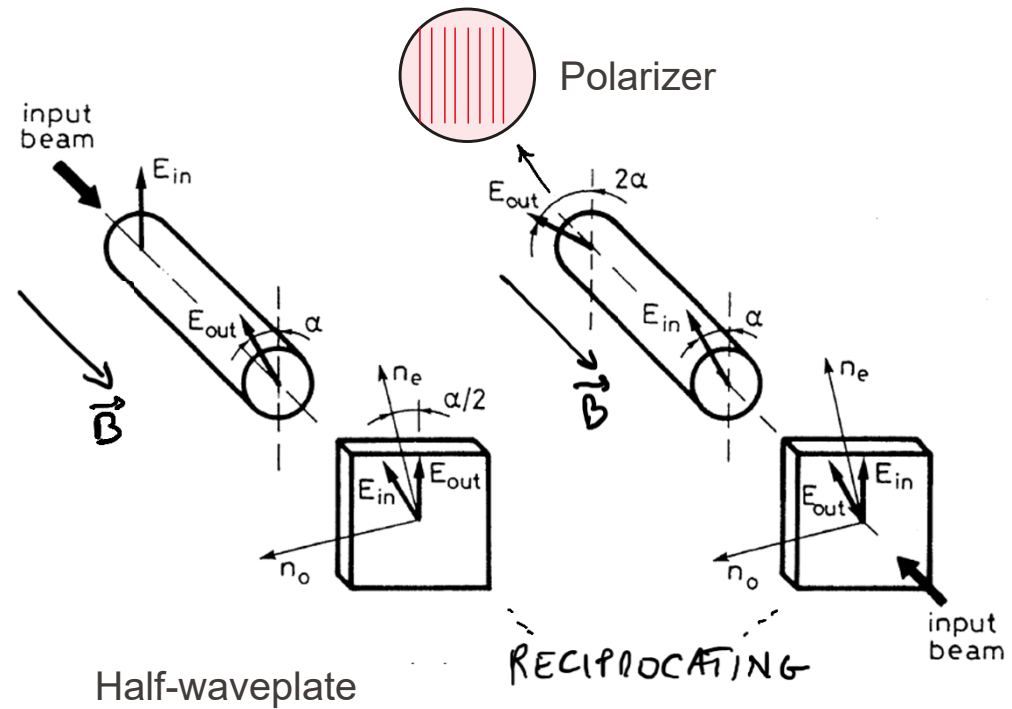
- How to prevent back-amplification?

Optical isolator



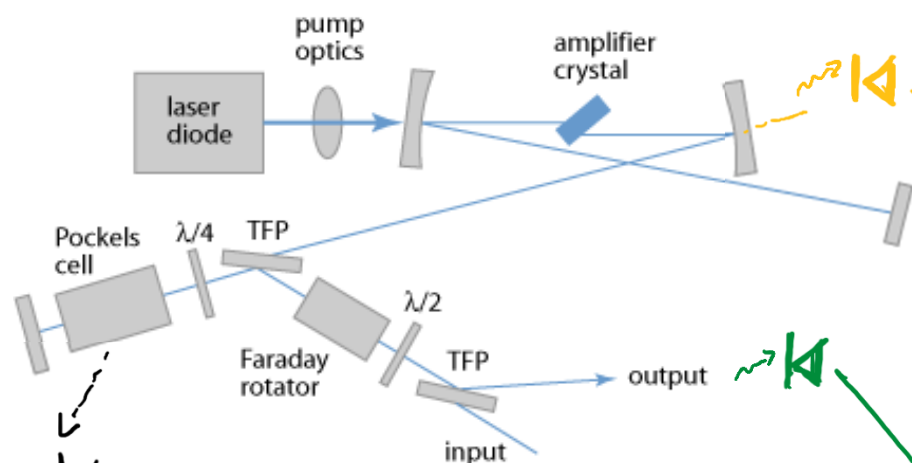
## Optical isolators

- Faraday effect
- Non-reciprocating optical element



Optical isolator is realized for  $\alpha=45^\circ$  by inserting a polarizer at the input

## Regenerative amplifier



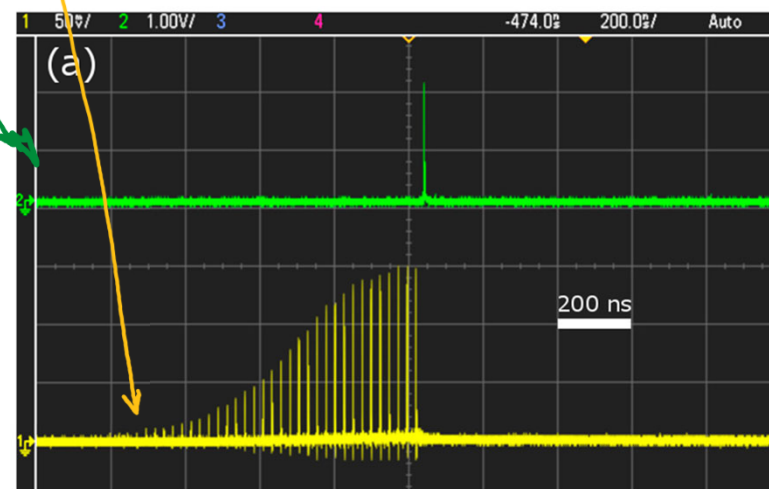
ON :  $\lambda/4$   
 OFF : no effect

Good efficiency, good mode quality.  
 kHz, mJ pulses at  $<100$  fs ( $> 10$  GW)  
 The Q-switch of the pump and the cell have to be well synchronized ( $<ns$ ).

The pulse train is incoupled in a cavity: only one pulse is selected by the switching of the the pockels cell

A large number of passes is then achieved, until the cell is switched back

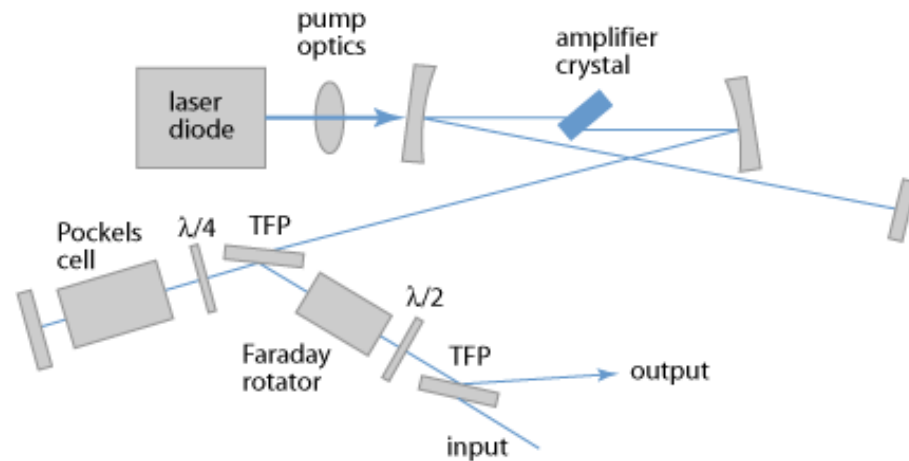
Approach the saturation intensity, efficiently depleting the gain medium



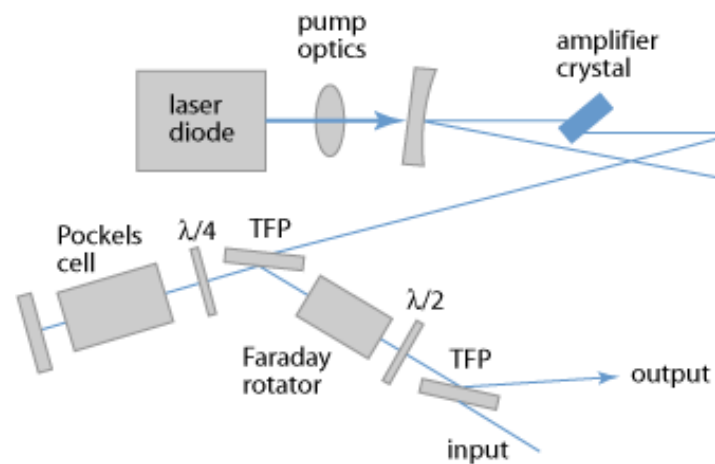
OFF

No polarization rotation

Light does 1 x roundtrip

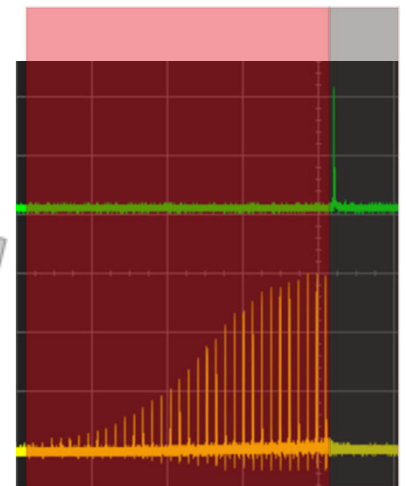


ON

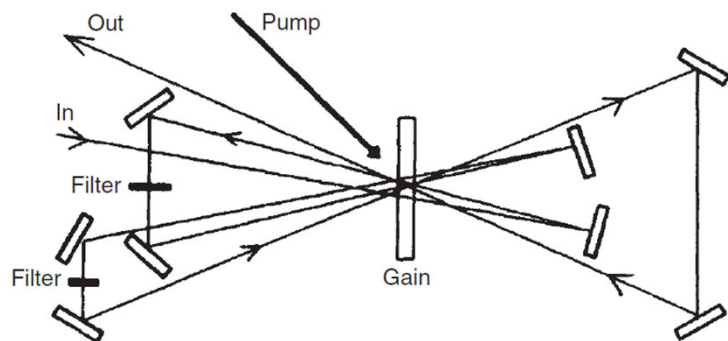
Voltage  $V = V_{\lambda/4}$ Equivalent to a  $\lambda/4$  plate

ON

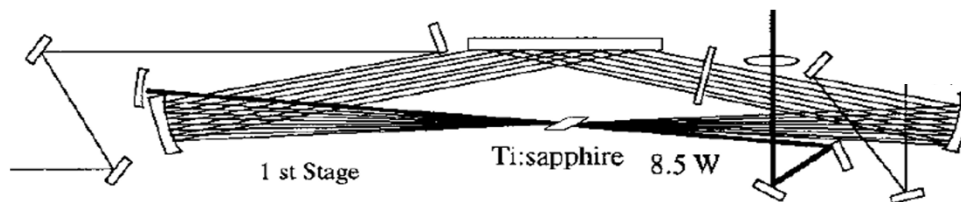
OFF



## Multi-pass amplifiers

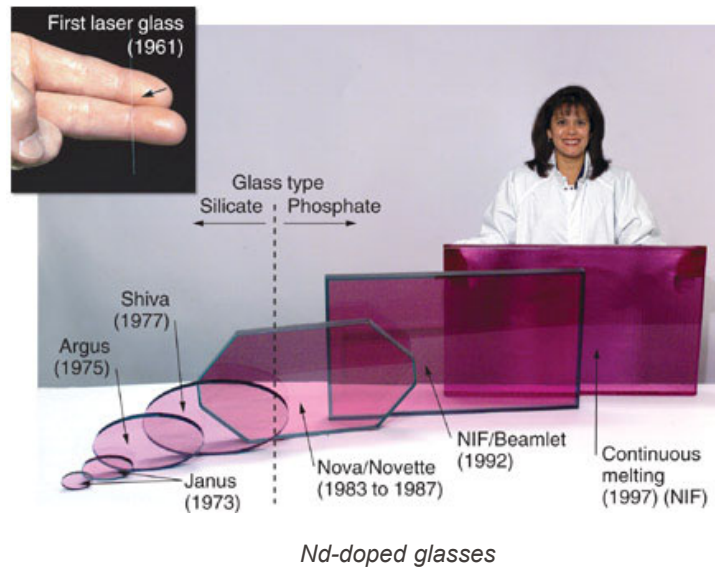


- Complex optical alignment
- Often used as “booster” stage of a reg. amplifier
- Flexibility between successive passes (change the mode size/apply filters )

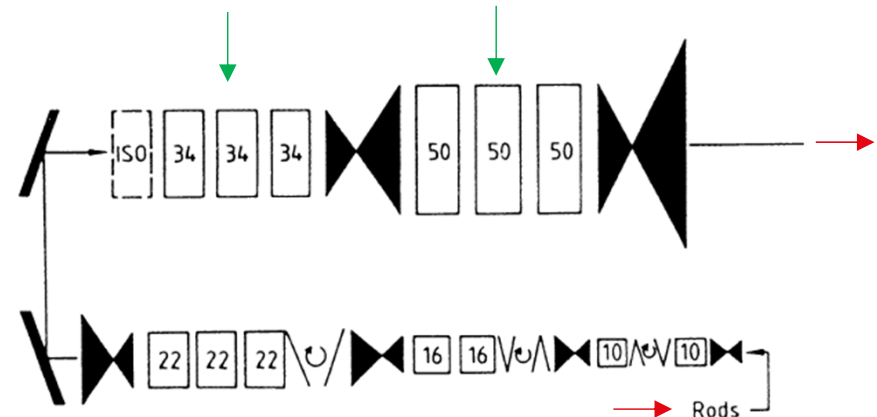


- Complex optical alignment

- In very high power system (>TW) regen. and multipass amplifiers are followed by amplification in low-gain, high aperture amplification stages



Flash-lamp pumped, very low rep. rate



- N N-cm aperture amplifiers
- Spatial filters
- Faraday isolators
- Mirrors
- Focus lens, window, debris shield
- Converter crystal array

Example: 1 amplification arm of the NOVA laser (LLNL)

10s kJ, ns pulses

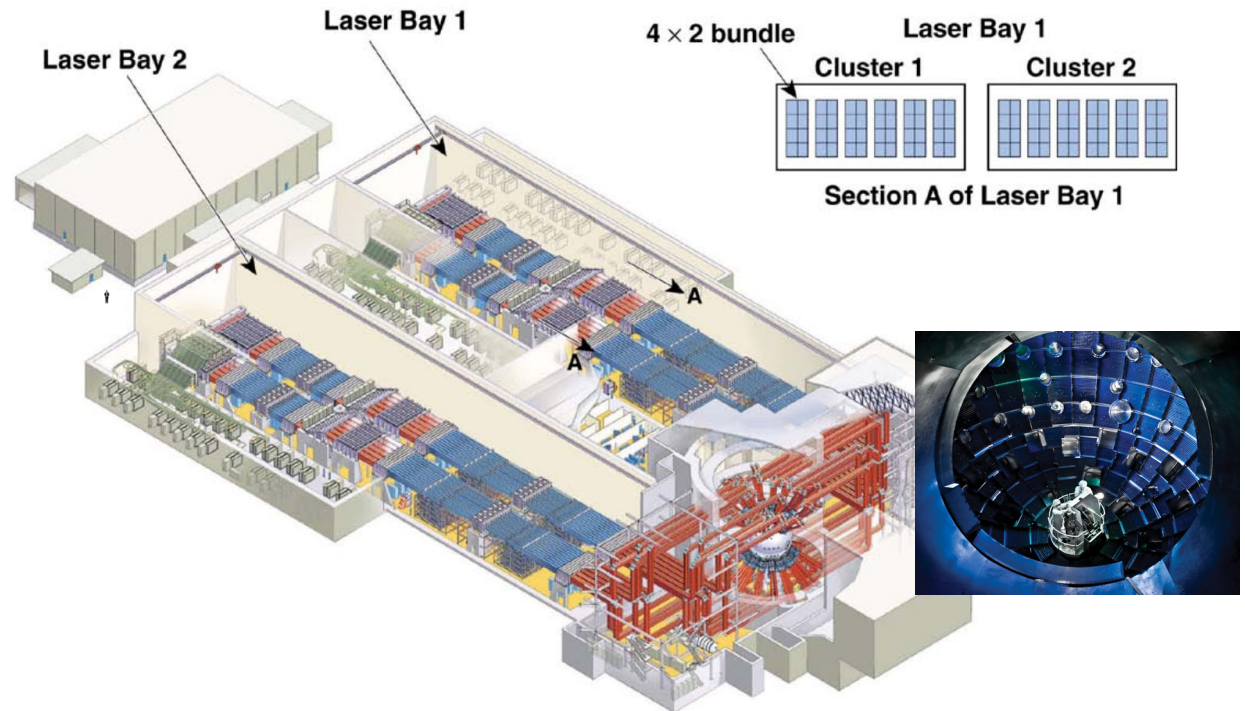
# The national ignition facility

Lasers for fusion research  
/plasma physics /high energy  
physics

1 x oscillator  
192 lasers Nd:glass  
amplifiers (20 kJ) in parallel

National Ignition Facility (USA)  
Laser Megajoule (France)

Pulse duration **20 ns**, 4.2 MJ  
(1  $\mu\text{m}$  wavelength); (1.8 MJ @  
0.34  $\mu\text{m}$ )



Hard to further scale the pulse power!

A decrease of the pulse duration is used to reach the **Petawatt level ( $10^{15}$  W)**:

Required the introduction of CPA – Chirped pulse amplification

VULCAN laser UK 1 PW,  $\approx 0.5$  ps

LFEX Japan : 2 PW,  $>1$  ps

ELI-NP (EU, planned) Extreme light infrastructure

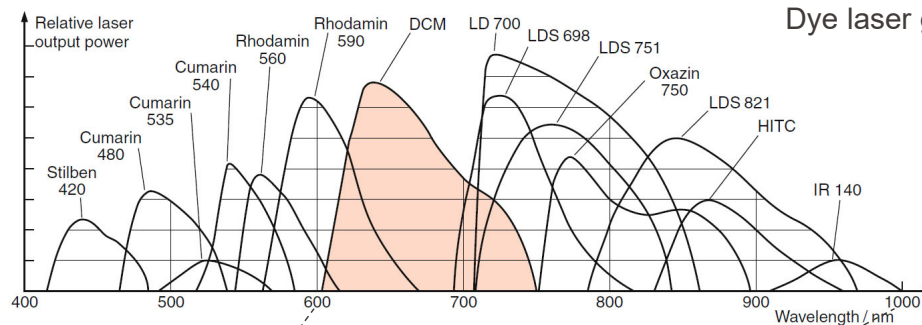
2 x 10 PW (with **<50 fs**; Ti:Sapphire amplifiers)

## Part 2 Amplifiers of short pulses

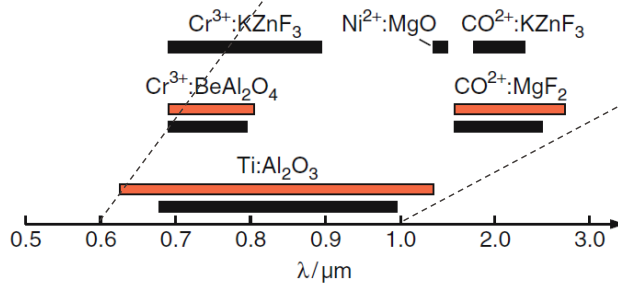
$$\Delta\nu \cdot \tau_p \geq k_{\text{BP}} \sim 1$$

## Time-bandwidth product

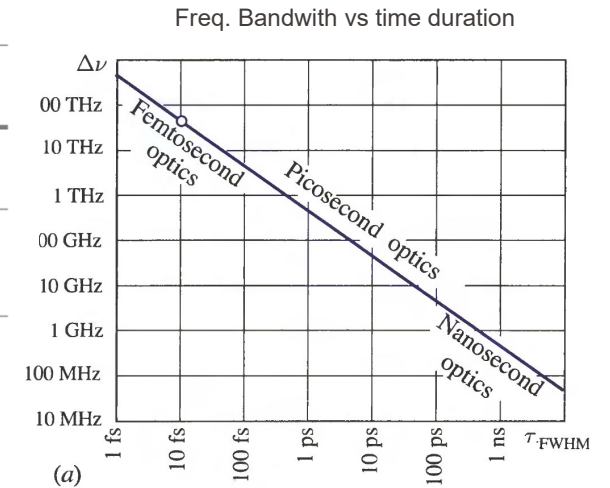
|                                   | $I(t) \ (x \equiv t/\tau)$ | $\tau_p/\tau$   | $\Delta\nu_p \times \tau_p$ |
|-----------------------------------|----------------------------|-----------------|-----------------------------|
| Gaussian                          | $I(t) = e^{-x^2}$          | $2\sqrt{\ln 2}$ | 0.4413                      |
| Hyperbolic secant (soliton pulse) | $I(t) = \text{sech}^2 x$   | 1.7627          | 0.3148                      |



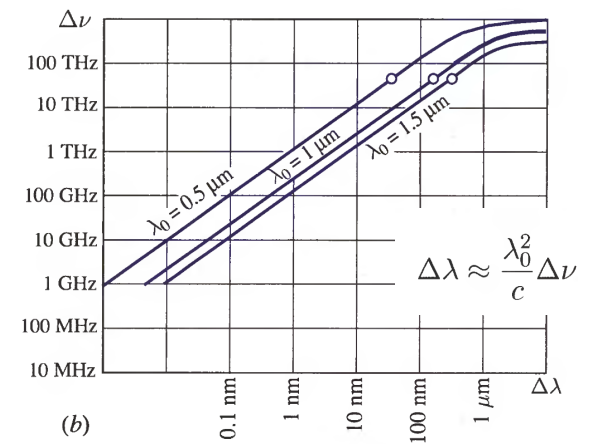
Dye laser gain



Vibronic lasers gain



(a)



(b)

Below 1 ps, we speak of ultrashort pulses: the spectral bandwidth is necessarily high!  
Ti:Sapphire solid state material with the largest known bandwidth (230 nm)

# $\chi^{(3)}$ effects: the nonlinear index of refraction

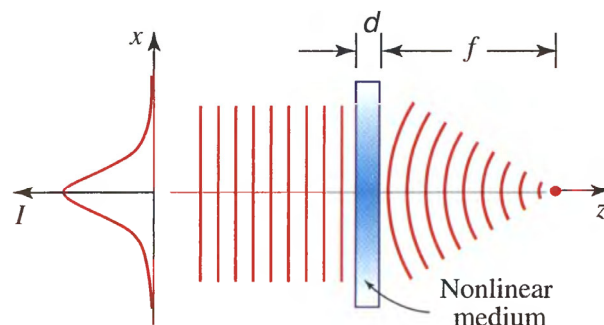
$$\tilde{P}^{(3)}(t) = \chi^{(3)} \tilde{E}(t)^3$$

$$n(\omega, I) = n_0(\omega) + n_2(\omega)I.$$

Note: this is one of the many terms of four-wave mixing, where three photon at frequency  $\omega$  mix to generate the same frequency ( $\omega + \omega - \omega \rightarrow \omega$ )

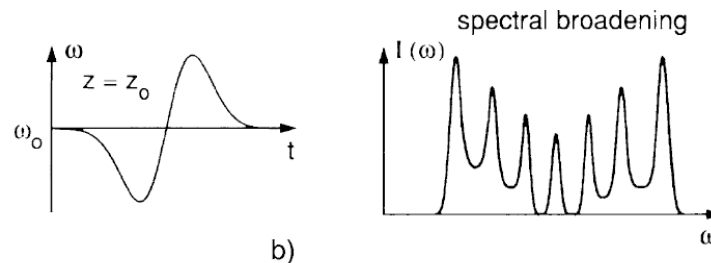
In space: Self-focusing

The radial variation of intensity translate into a radially varying optical path: this acts as a lens!



In time: Self-phase modulation and spectral broadening

$$\omega = \frac{d\phi}{dt} = \omega_0 - A \frac{dI}{dt}$$



## B integral

- The peak intensity during amplification increases as the pulse energy grows
- How to quantify the impact of  $n_2$  ?

$$B = \frac{2\pi}{\lambda} \int_0^L \underbrace{n_2}_{\substack{\text{NONLINEAR} \\ \text{INDEX OF REFRACTION}}} I(z) dz$$

Remember,  $\chi^{(3)}$  in time = self-phase modulation.

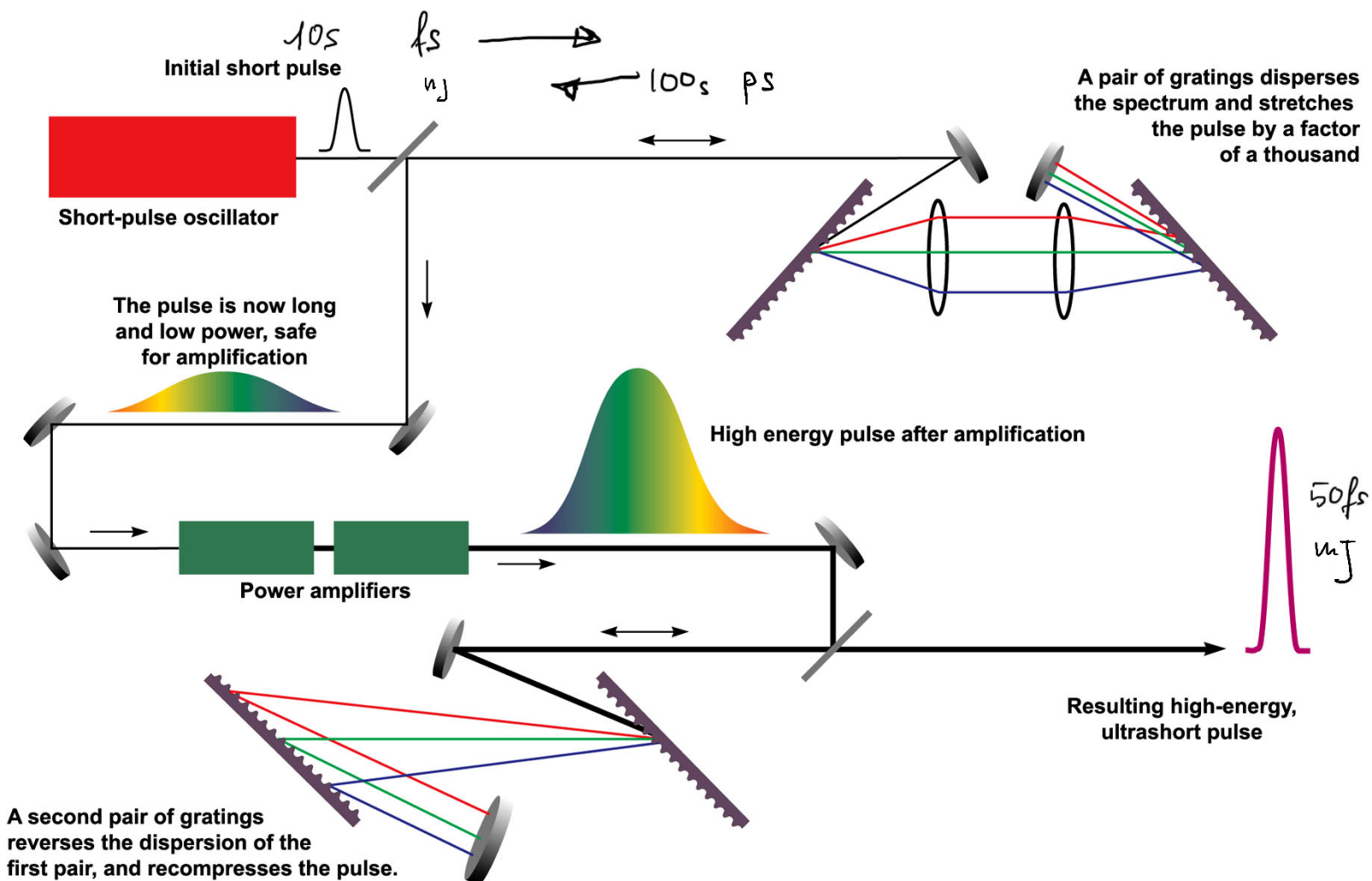
B-integral is a measure of the overall nonlinear phase accumulated. If B is too high, it is almost impossible to recompress the pulse.

Self-focusing might lead to damage of the optics

1. To further improve the laser intensities, the best route is to shorten the pulse duration
2. Self-effects become a fundamental limit when amplifying short ( picosecond and femtosecond) pulses

## CHIRPED PULSE AMPLIFICATION

# Chirped pulse amplification

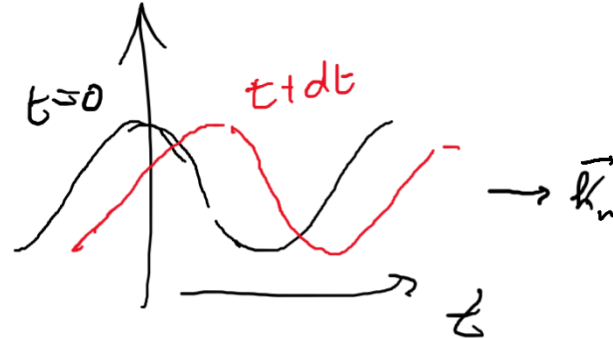


Pulse propagation some concepts:

$$E(z,t) = A \exp(i\phi(z,t)) = A \exp(i(\omega t - k_n z)) \quad \xrightarrow{\quad} k_n(\omega)$$

phase front  $d\phi = 0 \quad \rightarrow \quad \omega dt - k_n dz = 0$

$$v_{ph} = \frac{dz}{dt} = \frac{\omega}{k_n}$$

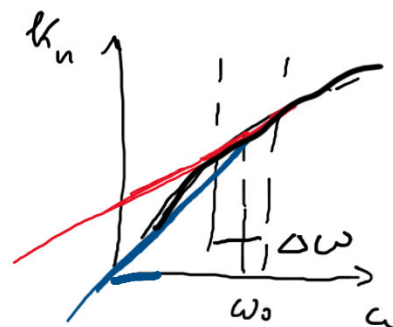


Phase velocity of plane wave in a material

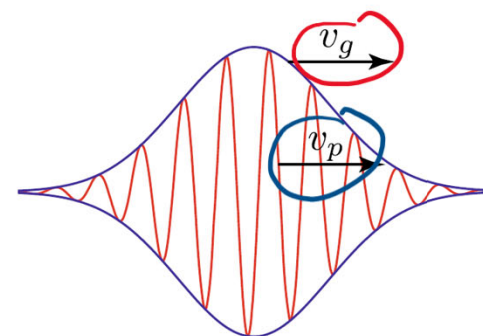
Wavepacket: 
$$E(t, z) = \int_{\Delta\omega} d\omega A_{\omega} e^{i(\omega t - k_n(\omega) z)} \longrightarrow E(t, 0) = A(t) e^{i\omega t}$$

For frequency narrow pulses (e.g. ps)  $\rightarrow$  small  $\Delta\omega$  around  $\omega_0$

$$k_n(\omega - \omega_0) = k_n(\omega_0) + k'_n(\omega_0) \Delta\omega$$



LINEAR APPROX



Phase vs Group velocity

DIFFERENT

$$E(t, z) \sim \underbrace{A(t - z/v_g)}_{\text{SAME SHAPE, TIME OFFSET}} \exp(i\omega t - k_n(\omega_0) z)$$

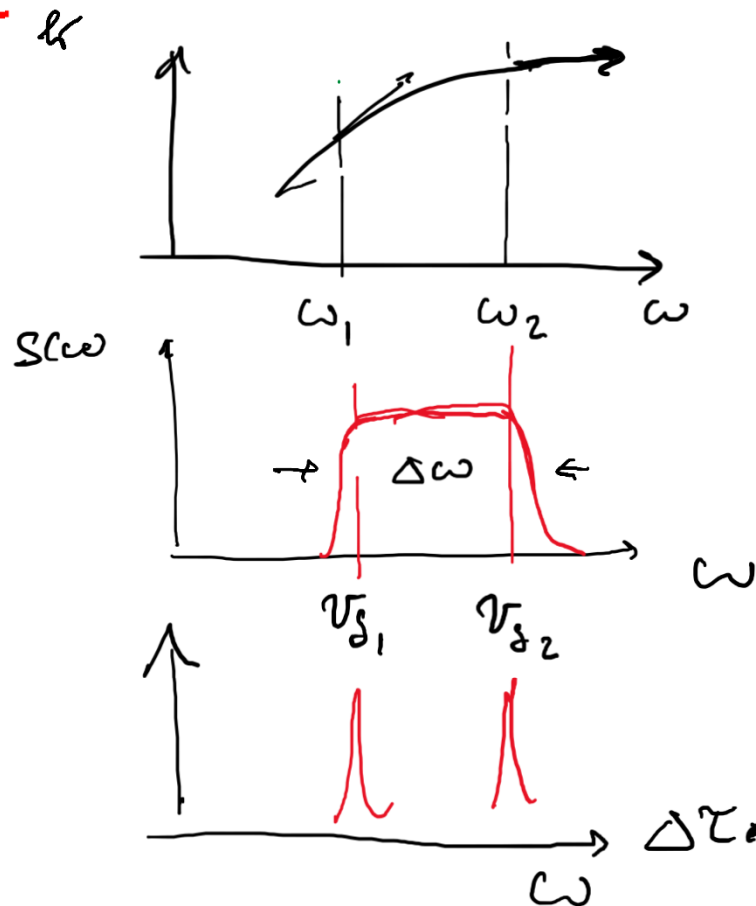
$$1/v_g = \frac{\partial k_n}{\partial \omega} \quad v_g = \frac{\partial \omega}{\partial k_n}$$

SPECTRAL  
PHASE

$$\phi(\omega - \omega_0) = k_n(\omega - \omega_0) z$$

$\phi'$  "GROUP DELAY"

## Broadband frequency pulse propagation:



BANDWIDTH IS LARGE :  
AS THE PULSE PROPAGATES  
DIFFERENT PORTIONS OF THE SPECTRUM  
HAVE DIFFERENT GROUP VELOCITIES,

→ ESTIMATE OF  $\tau$  AFTER PROP. ?

$$\Delta\tau \sim L \Delta v_g^{-1} =$$

$$\phi'(\omega_2) - \phi'(\omega_1) \simeq \underbrace{\phi''(\omega_0)}_{\text{GROUP DELAY DISPERSION}} (\omega_2 - \omega_1)$$

$$\Delta\tau \sim \phi''(\omega_0) \Delta\omega$$

$$L \left| \frac{d^2 k_n}{d\omega^2} \right|_{\omega=\omega_0} \Delta\omega$$

GROUP  
VELOCITY  
DISPERSION

GROUP DELAY  
DISPERSION  
 $\left. \frac{d}{d\omega} \left( \frac{1}{v_g} \right) \right|_{\omega=\omega_0}$

General idea: - expand the phase of the pulse around the central frequency  
- calculate the temporal shape by inverse Fourier transformation

REAL E-FIELD

TIME DEP. PHASE

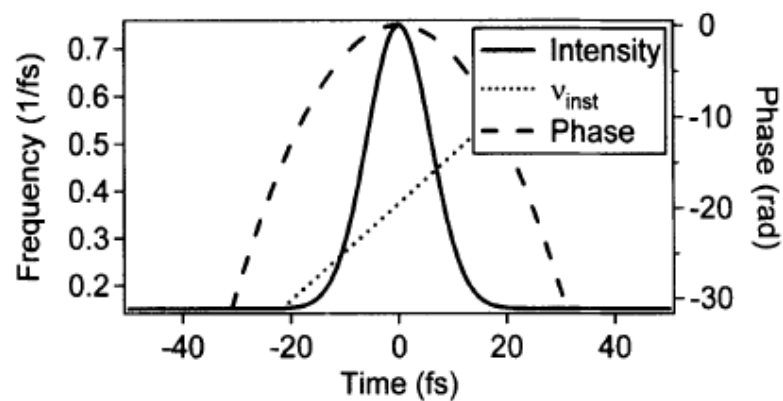
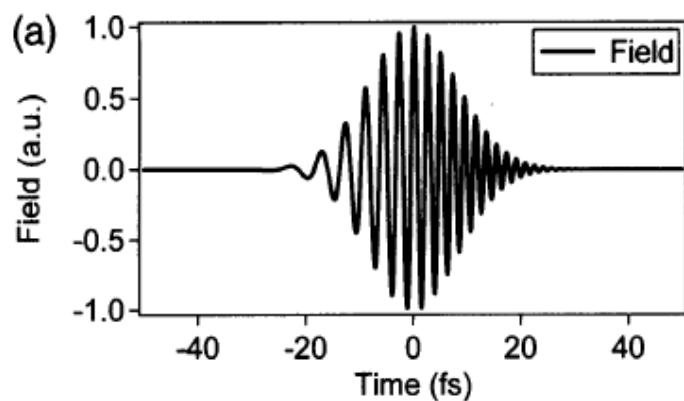
$$\mathcal{E}(t) = \frac{1}{2} \sqrt{I(t)} \exp\{i [\omega t - \phi(t)]\} + c.c.$$

FAST OSCILLATING CARRIER

COMPLEX AMPLITUDE (NO CARRIER)

$$E(t) \equiv \sqrt{I(t)} \exp[-i\phi(t)]$$

Pulse duration:  
FWHM  $I(t)$



$$\omega_{\text{inst}}(t) \equiv \omega_0 - d\phi/dt$$

$\phi(t)$  is the TEMPORAL PHASE: information of frequency vs time

## TIME vs FREQUENCY DOMAIN

$$\mathcal{E}(t) = \frac{1}{2} \sqrt{I(t)} \exp\{i[\omega t - \phi(t)]\} + c.c.$$

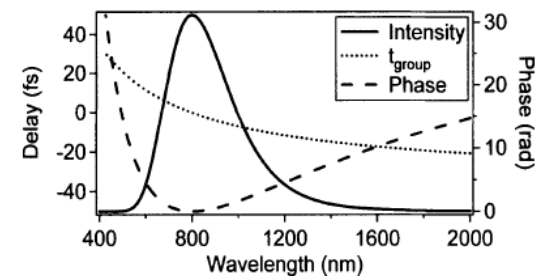
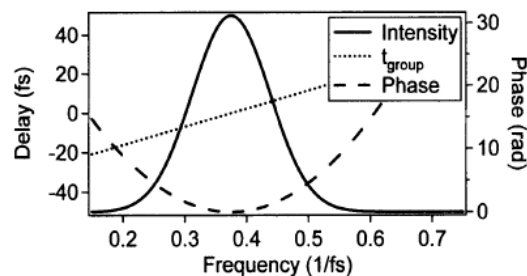
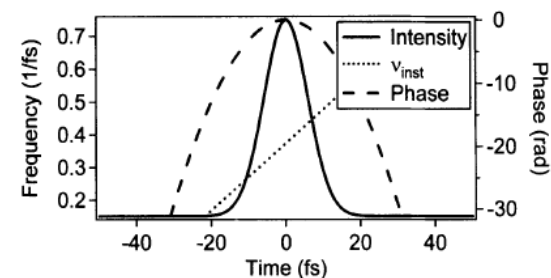
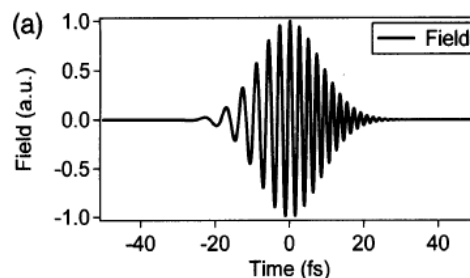
$$\tilde{\mathcal{E}}(\omega) = \int_{-\infty}^{\infty} \mathcal{E}(t) \exp(-i\omega t) dt$$

FOURIER TRANSFORM

Spectral description:

$$\tilde{\mathcal{E}}(\omega) = \sqrt{S(\omega)} \exp[-i\varphi(\omega)]$$

COMPLEX  
AMPLITUDE



$\varphi(\omega)$  is the SPECTRAL PHASE:  
information of time vs frequency

# Analytical example : Gaussian pulse with complex parameter

$$E(t) = \exp[-\Gamma t^2] \exp[i\omega_0 t]$$

$$\Gamma = \Gamma_1 - i\Gamma_2$$

L.I.D. CHANGE  
OF FREQ. WITH  
TIME

$$\phi(t) = \omega_0 t + \Gamma_2 t^2 \Rightarrow \dot{\phi}(t) = \omega_0 + \Gamma_2 t = \omega(t)$$

$$I(t) = |E|^2 \quad \tau_p = \sqrt{\frac{2 \ln(2)}{\Gamma_1}} \quad \longleftrightarrow \quad g(t) = \exp\left[-4 \ln 2 \left(\frac{t}{\tau_p}\right)^2\right]$$

(FWHM)

$$\tilde{E}(\omega) = \mathcal{F}[E(t)] \propto \exp\left[-\frac{(\omega - \omega_0)^2}{4\Gamma}\right]$$

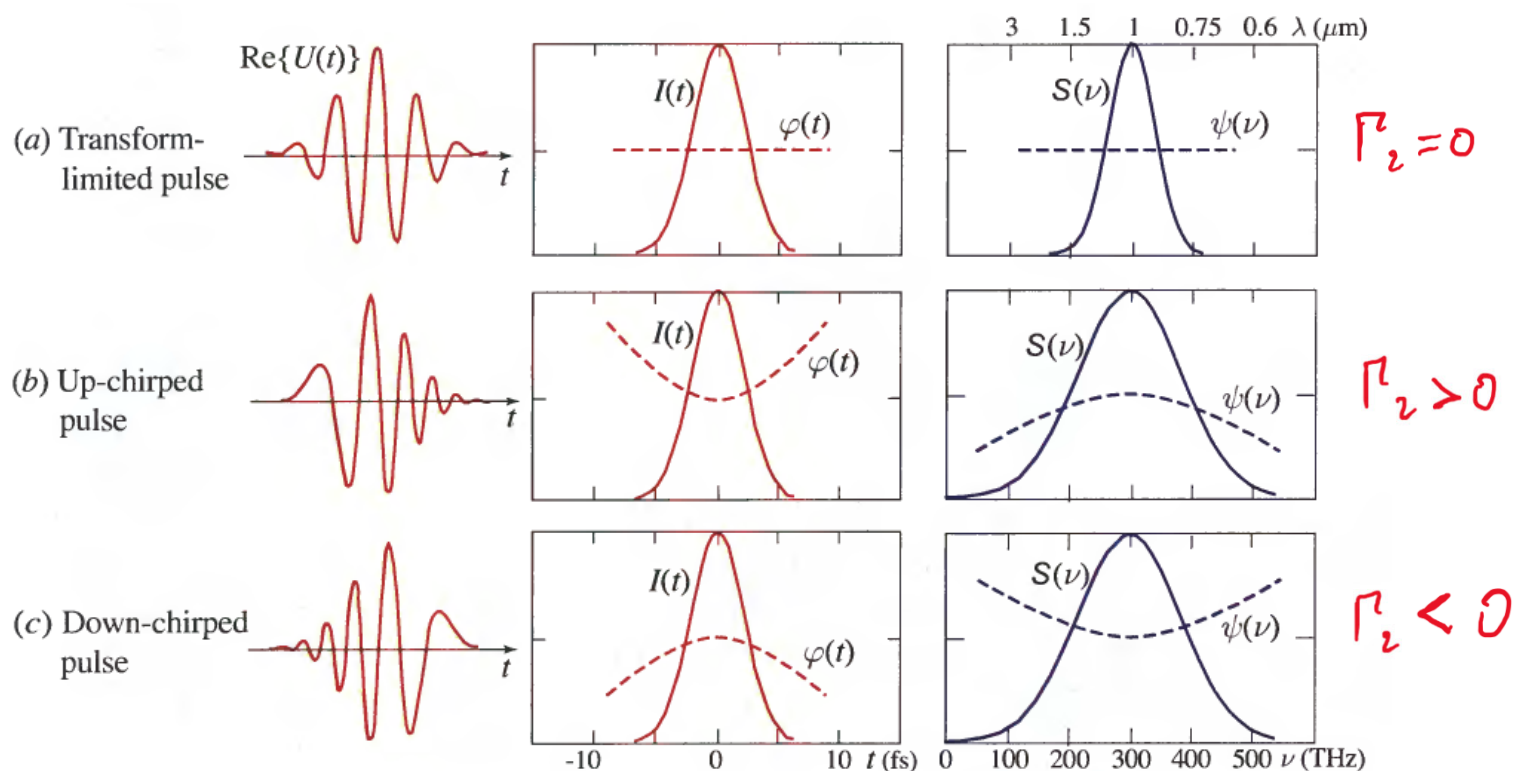
$\Gamma_2 \rightarrow 0$  Fourier limit  
 $\Gamma_2 \neq 0$  chirped pulse

$$\Delta\nu_p = \frac{\Delta\omega}{2\pi} = \frac{\sqrt{2 \ln 2}}{\pi} \sqrt{\Gamma_1 \left(1 + \left(\frac{\Gamma_2}{\Gamma_1}\right)^2\right)}$$

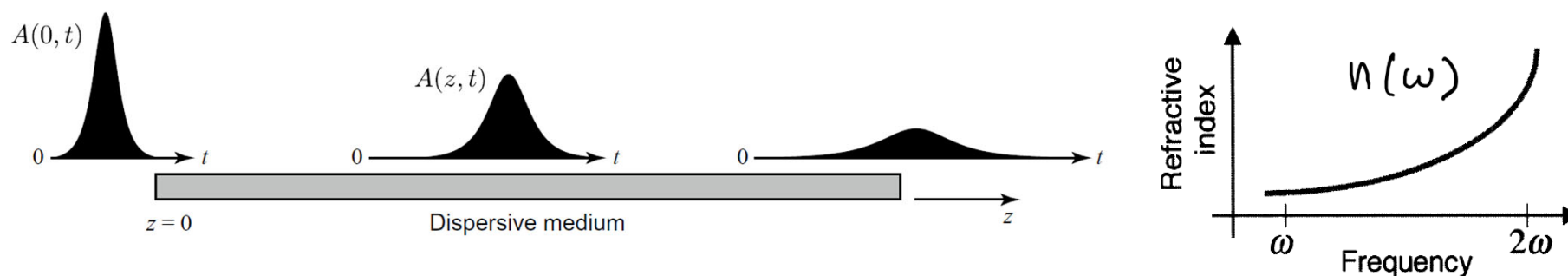
SPECTRAL BROAD. *min*

$$\tau_p \Delta\nu_p = \frac{2 \ln 2}{\pi} \sqrt{1 + \left(\frac{\Gamma_2}{\Gamma_1}\right)^2}$$

# Chirped pulses



- A non-flat phase is the result of a non-transform limited pulse: by compensating the phase a shorted pulse could be produced.
- By controlling the phase one can stretch/compress the pulse.

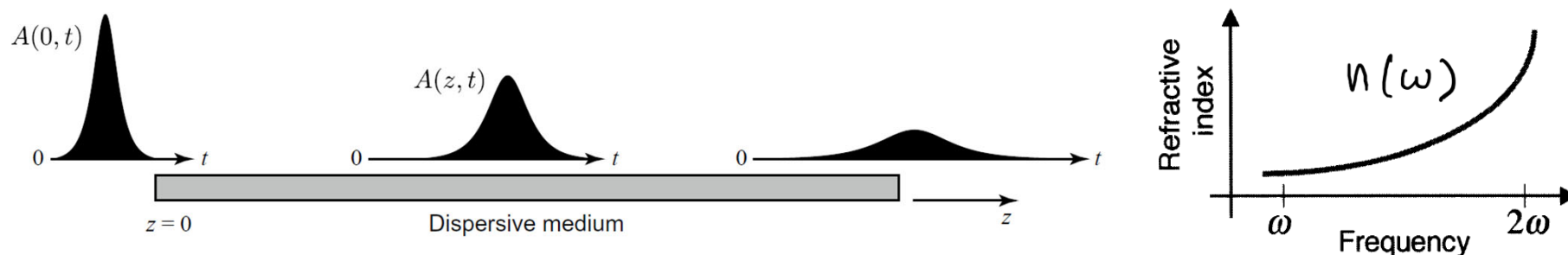


- **linear** medium: no nonlinear effects (e.g. gain or saturable absorption)
- Spectral intensity envelope is unchanged after propagation
- SVEA : Slowly Varying Envelope Approximation

SIMPLIFIED  
WAVE  
EQUATION.

$$\left| \frac{\partial^2 \tilde{A}}{\partial t^2} \right| \ll \left| k_n(\omega) \frac{\partial \tilde{A}}{\partial z} \right|$$

ENVELOPE DOES NOT  
CHANGE MUCH OVER  
A DISTANCE  
 $dz \sim \lambda_n$



WAVE EQ.  
IN THE  
FREQ. DOMAIN

$$\tilde{A}(z, \omega) = \tilde{A}(0, \omega) \exp[i\varphi(\omega)]$$

$$\varphi(\omega) = k(\omega)z = \frac{\omega n(\omega) z}{c}$$

TAYLOR EXPANSION  
OF THE SPECTRAL  
PHASE

$$\varphi(\omega) = \varphi_0 + (\omega - \omega_0) \varphi_1 + (\omega - \omega_0)^2 \varphi_2 / 2 + \dots$$

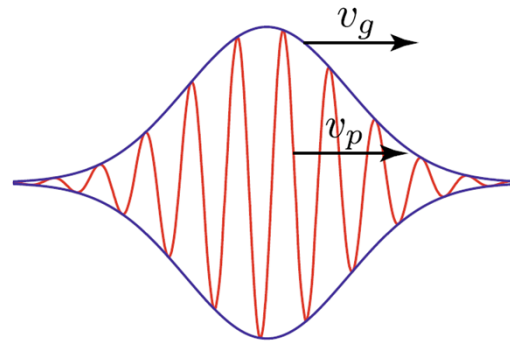
$$\downarrow \quad \downarrow \quad \downarrow$$

$$\frac{\partial \varphi}{\partial \omega} \text{ GD} \quad \frac{\partial^2 \varphi}{\partial \omega^2} \text{ GDD} \quad \dots \quad \frac{\partial^3 \varphi}{\partial \omega^3} \text{ TOD}$$

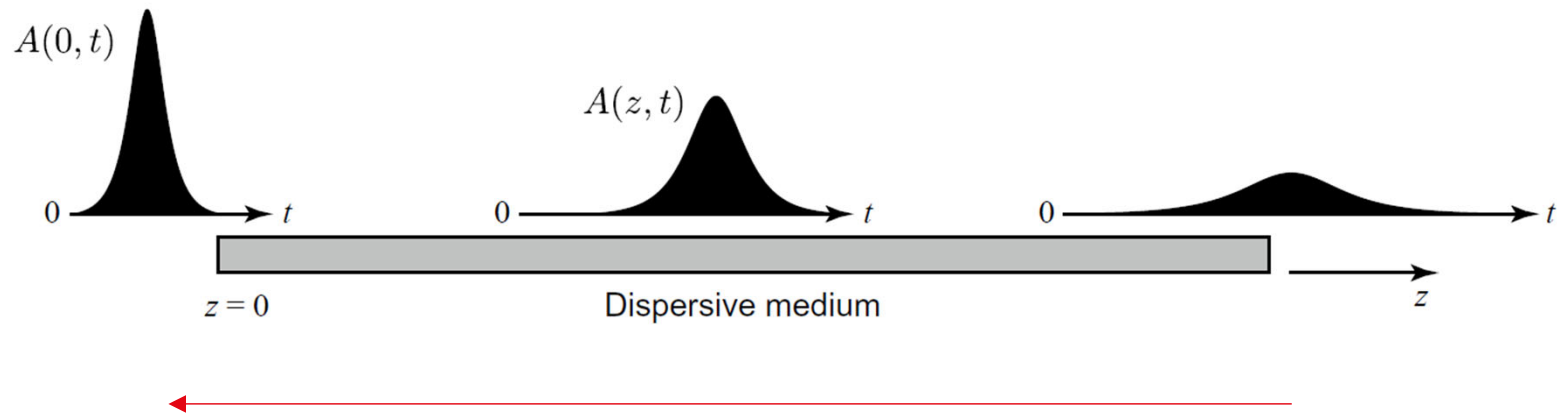
FOR A GAUSSIAN PULSE  
 → ANALYTICAL SOLUTIONS  
 (up To 2<sup>nd</sup> order dispersion)

$$\frac{1}{\Gamma(L_d)} = \frac{1}{\Gamma(0)} + 2ik_n'' L_d$$

$$E(L_d, t) \propto \exp\left[i\omega_0 \left(t - \frac{L_d}{v_p(\omega_0)}\right)\right] \exp\left[-\Gamma(L_d) \left(t - \frac{L_d}{v_g(\omega_0)}\right)^2\right]$$

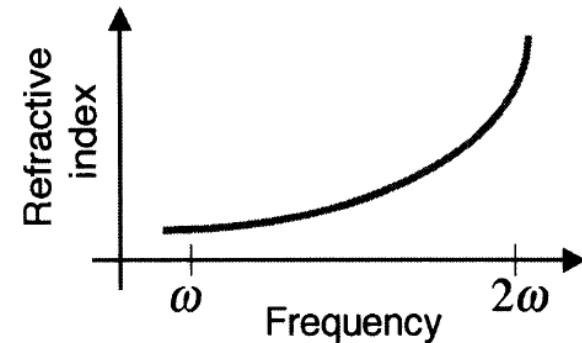


| Dispersion parameter                                               | Definition                                                           |
|--------------------------------------------------------------------|----------------------------------------------------------------------|
| Phase velocity $v_p$                                               | $\omega/k_n$                                                         |
| Group velocity $v_g$                                               | $\frac{d\omega}{dk_n} = \left(\frac{dk_n}{d\omega}\right)^{-1}$      |
| Group delay $T_g$<br>(first order dispersion)                      | $\frac{d\phi}{d\omega} = L_d \frac{dk_n}{d\omega} = \frac{L_d}{v_g}$ |
| Group delay dispersion $dT_g/d\omega$<br>(second order dispersion) | $\frac{d^2\phi}{d\omega^2}$                                          |
| $\frac{dT_g}{d\lambda}$                                            | $d\lambda = -\frac{\lambda^2}{2\pi c} d\omega$                       |
| Third order dispersion                                             | $\frac{d^3\phi}{d\omega^3}$                                          |

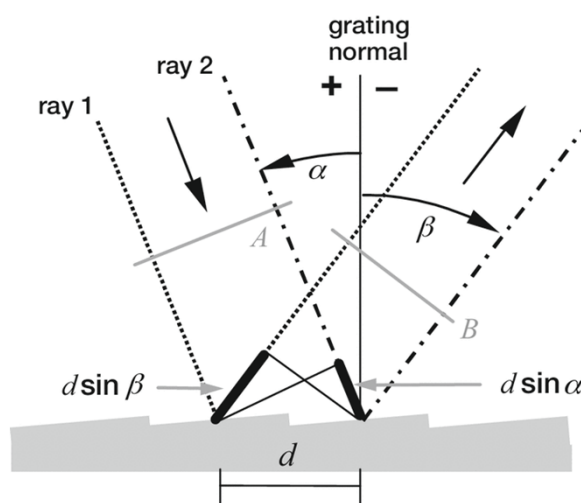


- Most material have “normal” dispersion at VIS-NIR wavelengths
- To compensate linear propagation temporal broadening, we need optical devices with opposite (“anomalous”) dispersion.
- Multiple materials in an amplifier, multiple passes.

Calculate the cumulative GDD (and TOD, and higher orders) -> Compensate with a suitable compressor (e.g. prism compressor, grating compressor.. )



## DIFFRACTION GRATING



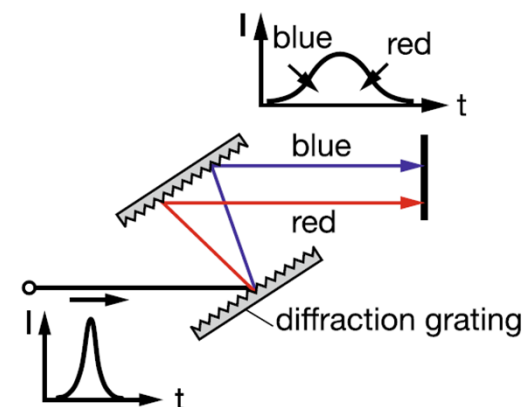
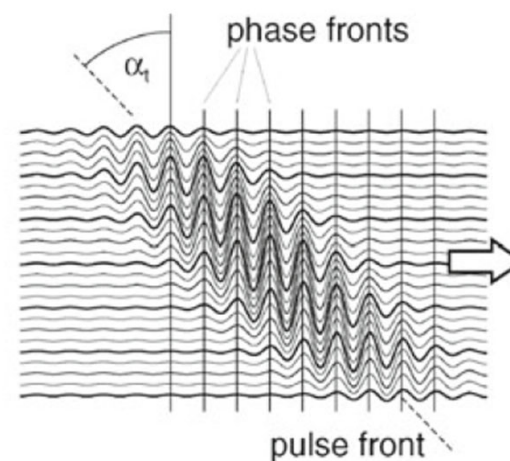
$$n\lambda = d(\sin \beta - \sin \alpha)$$

$$\beta = \sin^{-1} \left[ \frac{n\lambda}{d} + \sin \alpha \right]$$

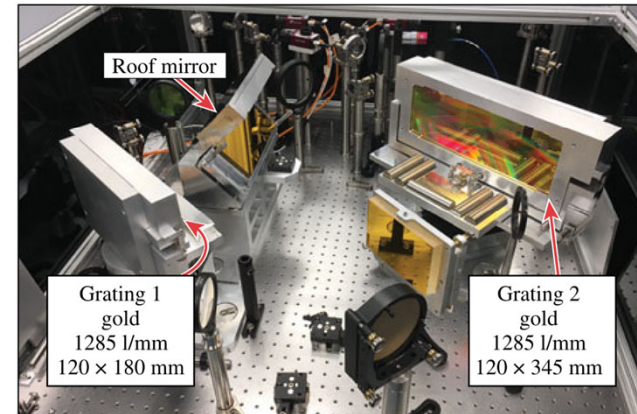
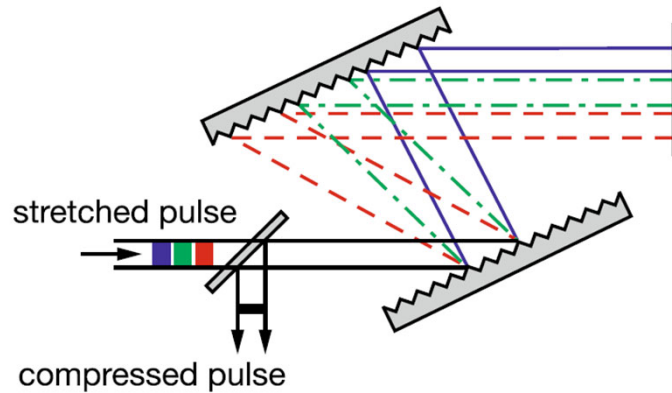
Periodic array of scatterers with spacing  $d = 1/\Lambda$

Angular dispersion: wavelength-dependent propagation angle.

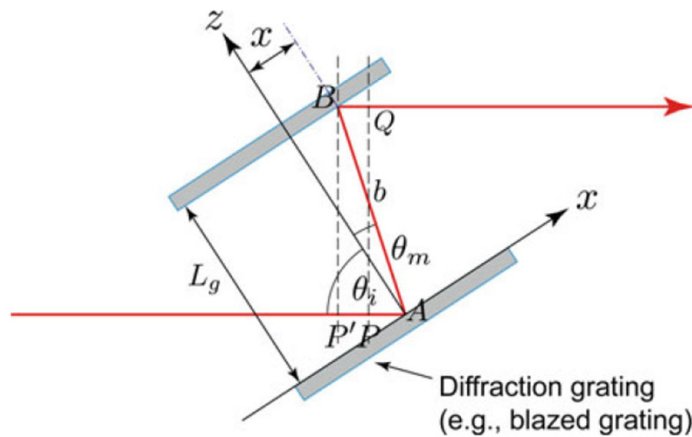
Pulse front tilt/spatial chirp: multiple grating configuration



## TREACY COMPRESSOR



- 4 gratings are necessary for avoiding pulse front tilt and spatial chirp
- High efficiency is fundamental : transmission  $\eta^4$
- Blazed gratings, near Littrow condition -> diffraction efficiency in the first order can approach 90%



2x contributions

$$\phi = \frac{\omega}{c}L + \phi_g(x)$$

$$x = L_g \tan \theta_m$$

1. Optical path

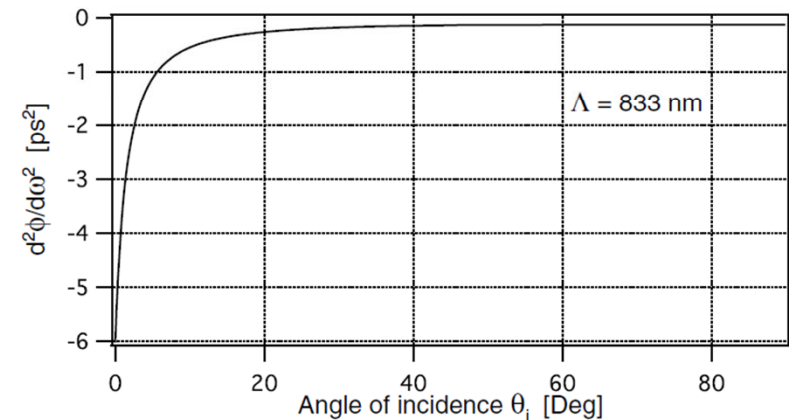
$$\frac{L_g}{\cos \theta_m} [1 + \cos (\theta_m + \theta_i)]$$

2. Spatial dependent phase shift due to diffraction

$$\phi_g(x) = \pi - m \frac{2\pi}{\Lambda} x$$

$$\frac{d^2\phi}{d\omega^2} = -\frac{\lambda^3 L_g}{\pi c^2 \Lambda^2} \left[ 1 - \left( \frac{\lambda}{\Lambda} - \sin \theta_i \right)^2 \right]^{-3/2}$$

|          |                                                                                                            |                                                                                                                                                                    |
|----------|------------------------------------------------------------------------------------------------------------|--------------------------------------------------------------------------------------------------------------------------------------------------------------------|
| Sapphire | $n(800 \text{ nm}) = 1.76019$                                                                              |                                                                                                                                                                    |
|          | $\left. \frac{\partial n}{\partial \lambda} \right _{800 \text{ nm}} = -0.0268 \frac{1}{\mu\text{m}}$      | $\left. \frac{\partial k_n}{\partial \omega} \right _{800 \text{ nm}} = 5.87 \times 10^{-9} \frac{\text{s}}{\text{m}}$                                             |
|          | $\left. \frac{\partial^2 n}{\partial \lambda^2} \right _{800 \text{ nm}} = 0.064 \frac{1}{\mu\text{m}^2}$  | $\left. \frac{\partial^2 k_n}{\partial \omega^2} \right _{800 \text{ nm}} = 5.80 \times 10^{-26} \frac{\text{s}^2}{\text{m}} = 58 \frac{\text{fs}^2}{\text{mm}}$   |
|          | $\left. \frac{\partial^3 n}{\partial \lambda^3} \right _{800 \text{ nm}} = -0.377 \frac{1}{\mu\text{m}^3}$ | $\left. \frac{\partial^3 k_n}{\partial \omega^3} \right _{800 \text{ nm}} = 4.21 \times 10^{-41} \frac{\text{s}^3}{\text{m}} = 42.1 \frac{\text{fs}^3}{\text{mm}}$ |



Example: Reg. Ti:Sapphire, 5 mm Crystal, 10 passes  $\rightarrow 58 \times 5 \times 10 = 2.9 \times 10^3 \text{ fs}^2$

Grating compressor, 1200 lines/mm,  $d = 5 \text{ cm}$ ,  $20^\circ$  incidence  $\rightarrow -0.2 \text{ ps}^2 = -0.2 \times 10^6 \text{ fs}^2$

- Gratings allow for a very large amount of GVD compared to propagation in materials
- In a compact setup from 10 fs  $\rightarrow$  500 ps (would require 50 m of fused silica)

Grating compressor  $\rightarrow$  Negative dispersion: pulse compression during amplification!

$$\frac{d^2\phi}{d\omega^2} = -\frac{\lambda^3 L_g}{\pi c^2 \Lambda^2} \left[ 1 - \left( \frac{\lambda}{\Lambda} - \sin \theta_i \right)^2 \right]^{-3/2}$$

$$\frac{d^2\phi}{d\omega^2} = -\frac{m^2 \lambda^3 M^2 L_{\text{eff}}}{2\pi c^2 \Lambda^2} \left[ 1 - \left( -m \frac{\lambda}{\Lambda} - \sin \theta_i \right)^2 \right]^{-3/2}$$

$L < f$  positive GDD

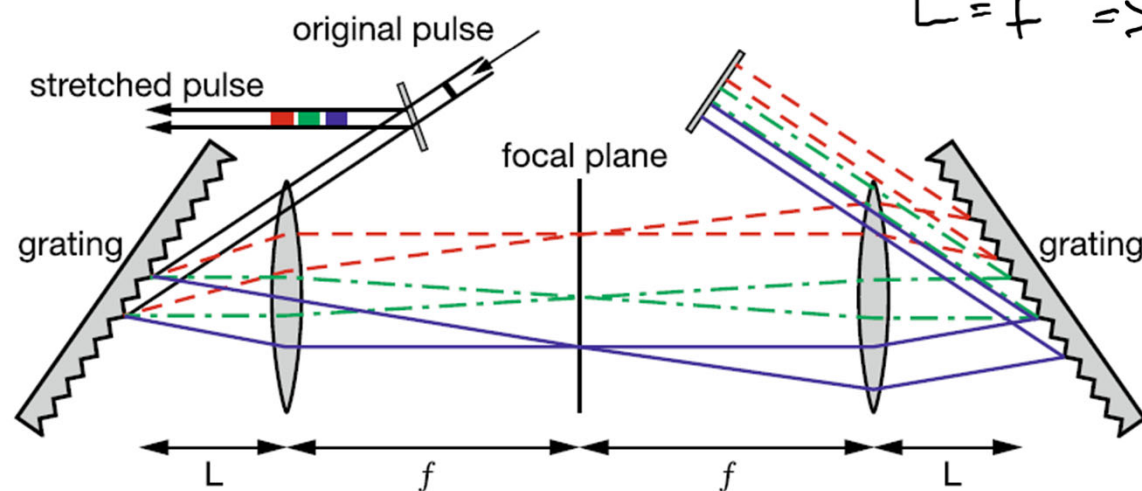
$L \geq f$  negative GDD

$L = f \Rightarrow 4f$  system

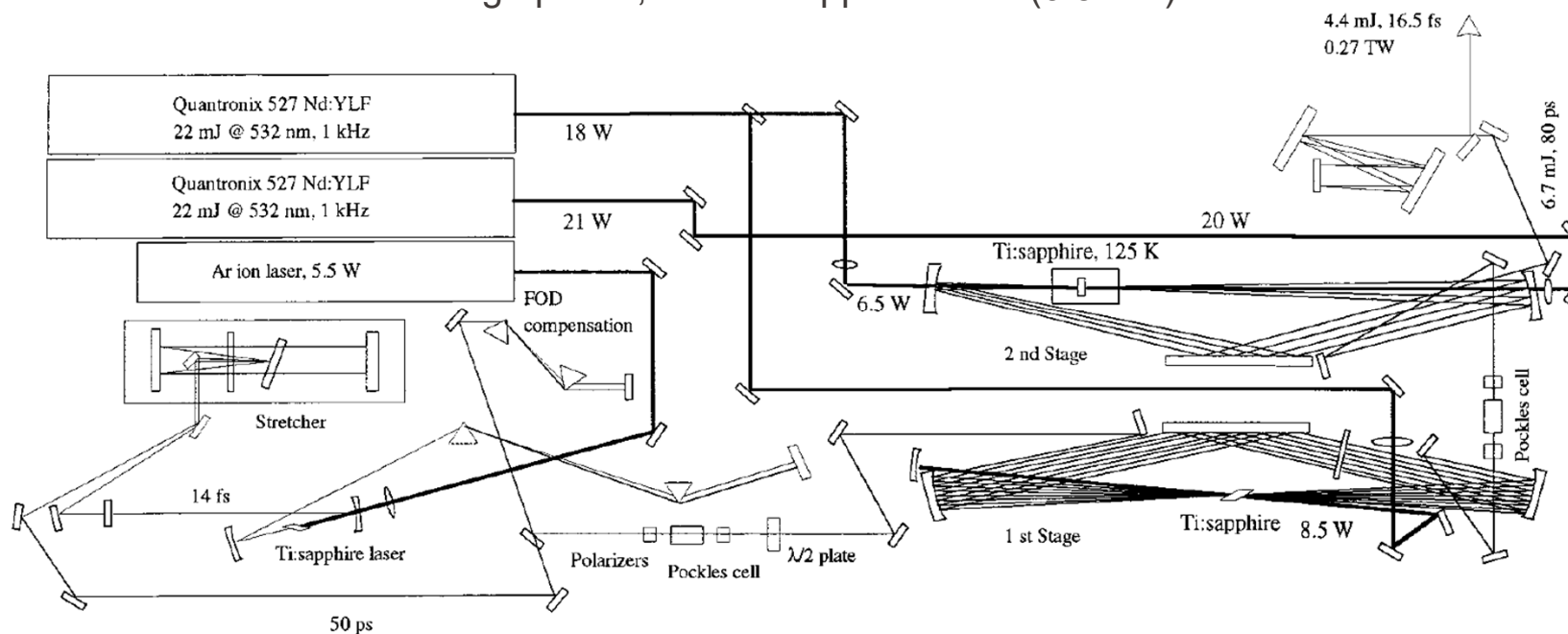
GDD = 0

CAN USE  
TO "SHAPE"  
THE PULSE

- Oeffner stretcher



## A high power, kHz Ti:Sapphire laser (0.3 TW)



Note: to produce short pulses (<50 fs) the higher order dispersion parameters have to be carefully compensated

Laboratory scale Ti:Sapphire amplifier : 1 kHz, 10-20 W, 30-50 fs (Cryogenically-cooled)

DIFFERENT TECHNOLOGIES ARE NEEDED TO INCREASE THESE PARAMETERS

## GAIN NARROWING IN TI:SAPPHIRE

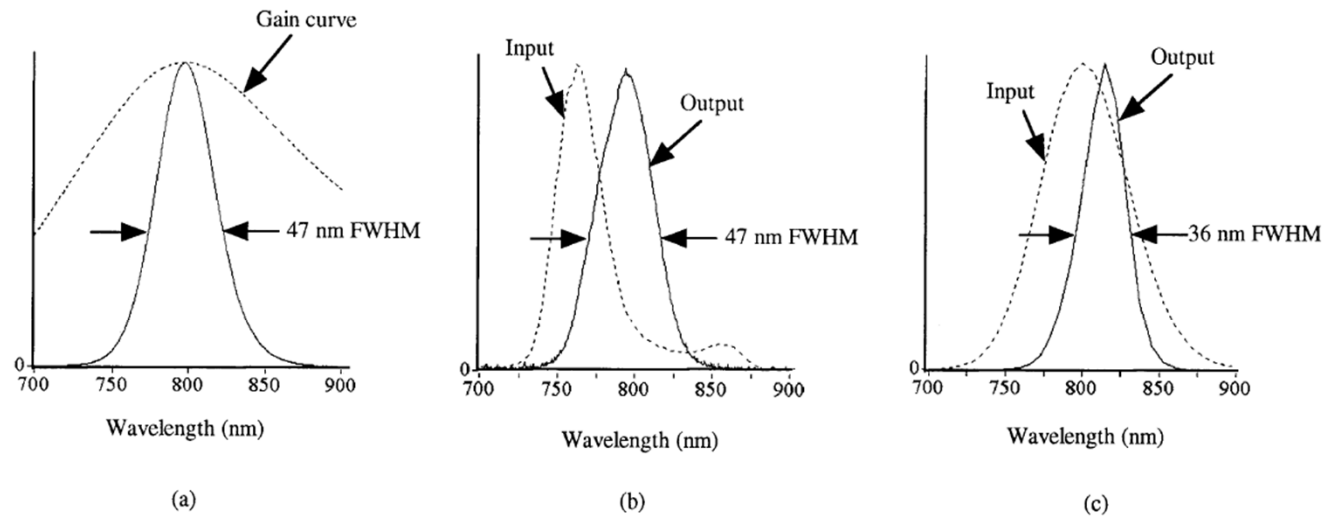


FIG. 5. Gain narrowing for the case of (a) an infinitely broad and flat input spectrum, (b) an optimally offset and shaped input spectrum, and (c) a nonoptimum input spectrum.

GAIN NARROWING IN TI:SAPPHIRE LIMITS THE  
ACHIEVABLE AMPLIFIED PULSE DURATION