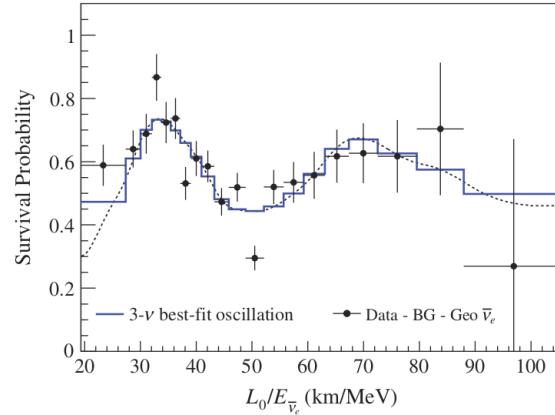


## Problem set 5: Solutions

### Problem 1

The KamLAND experiment measured the disappearance of  $\bar{\nu}_e$  anti-neutrinos produced in reactors located at  $\sim 180$  km from the detector. Estimate the parameter  $\Delta m_{12}^2$  from the plot of the survival probability as a function of  $L/E$ . Compare your result with the value determined by the KamLAND experiment:  $\Delta m_{12}^2 = (7.53 \pm 0.18) \times 10^{-5} \text{ eV}^2$ .



#### Solution:

The oscillation probability  $P_{\text{osc}}$  is given by the formula

$$P_{\text{osc}} = \sin^2 2\theta \sin^2 \left( \frac{1.27 \Delta m^2 L}{E} \right),$$

where the oscillation period is  $\pi$ . Then, considering the equality

$$\frac{1.27 \Delta m^2 L}{E} = \pi,$$

we obtain

$$\Delta m^2 = \frac{\pi}{1.27 (L/E)_{\text{period}}},$$

In order to estimate the period, we measure the separation between the first minimum and the second maximum on the KamLAND plot and multiply it by two. We estimate that the uncertainty is given by the bin width. This leads to  $(L/E)_{\text{period}} \simeq (40 \pm 4) \text{ km MeV}^{-1}$ . Then:

$$\Delta m^2 = (6.2 \pm 0.6) \times 10^{-5} \text{ eV}^2.$$

This result is almost compatible with the value stated by KamLAND. On the plot from KamLAND, it is clear that the fit function is not a pure sine squared, as it is distorted by the distribution of distances to reactors and the energy spectrum. The oscillation period can therefore be more precisely measured by accounting for these effects.

## Problem 2

Show that the mass-squared eigenvalues for the hamiltonian describing neutrino oscillations in matter, expressed in the  $(\nu_e, \nu_\mu)$  basis, are

$$m^2 = \frac{1}{2}(\mu^2 + B) \pm \frac{1}{2}\sqrt{(\Delta m^2 \cos 2\theta - B)^2 + (\Delta m^2 \sin 2\theta)^2},$$

where  $\Delta m^2 = m_2^2 - m_1^2$ , and that the corresponding mixing angle is given by

$$\tan 2\theta_{\text{matter}} = \frac{\sin 2\theta}{\cos 2\theta - \frac{B}{\Delta m^2}},$$

where  $B = 2EV_W$  accounts for the charged-current interaction potential.

### Solution:

In order to find the mass-squared eigenvalues for the hamiltonian describing neutrino oscillations in matter, we diagonalise the matrix  $H_\nu$ , containing an additional potential which is diagonal in the  $(\nu_e, \nu_\mu)$  basis. Then the matrix becomes

$$\frac{1}{2E} \begin{pmatrix} M_{ee}^2 + B & M_{\mu e}^2 \\ M_{\mu e}^2 & M_{\mu \mu}^2 \end{pmatrix},$$

where we use the following expressions for the matrix elements

$$\begin{aligned} M_{ee}^2 &= \frac{1}{2}(\mu^2 - \Delta m^2 \cos 2\theta), \\ M_{\mu \mu}^2 &= \frac{1}{2}(\mu^2 + \Delta m^2 \cos 2\theta), \\ M_{e \mu}^2 &= \frac{1}{2}\Delta m^2 \sin 2\theta. \end{aligned}$$

Solving the eigenvalue problem

$$\left[ \frac{1}{2}(\mu^2 - \Delta m^2 \cos 2\theta) + B - m^2 \right] \left[ \frac{1}{2}(\mu^2 + \Delta m^2 \cos 2\theta) - m^2 \right] - \frac{1}{4}(\Delta m^2 \sin 2\theta)^2 = 0,$$

we find

$$m^2 = \frac{\mu^2 + B}{2} \pm \frac{\Delta m^2}{2} \sqrt{\left( \cos 2\theta - \frac{B}{\Delta m^2} \right)^2 + \sin^2 2\theta} = \frac{1}{2}(\mu^2 + B \pm \Delta m_{\text{matter}}^2). \quad (1)$$

We see that in the limit  $B \rightarrow 0$  (vacuum), Eq.1 can be rewritten as

$$m_{B \rightarrow 0}^2 = \frac{\mu^2}{2} \pm \frac{\Delta m^2}{2} \sqrt{\cos^2 2\theta + \sin^2 2\theta} = \frac{1}{2}(\mu^2 \pm \Delta m^2). \quad (2)$$

By comparing Eqs.1 and 2, we can write:

$$\Delta m_{matter}^2 = \Delta m^2 \sqrt{\left(\cos 2\theta - \frac{B}{\Delta m^2}\right)^2 + \sin^2 2\theta}.$$

Replacing this expression in Eq.1, we get:

$$m^2 = \frac{\mu^2 + B}{2} \pm \frac{\Delta m_{matter}^2}{2} \sqrt{\frac{\left(\cos 2\theta - \frac{B}{\Delta m^2}\right)^2}{\left(\cos 2\theta - \frac{B}{\Delta m^2}\right)^2 + \sin^2 2\theta} + \frac{\sin^2 2\theta}{\left(\cos 2\theta - \frac{B}{\Delta m^2}\right)^2 + \sin^2 2\theta}}. \quad (3)$$

Now, if we compare Eq. 3 and the similar expression in vacuum (Eq. 2), we can write:

$$m^2 = \frac{\mu^2 + B}{2} \pm \frac{\Delta m_{matter}^2}{2} \sqrt{\sin^2 2\theta_m + \cos^2 2\theta_m}, \quad (4)$$

where  $\theta_m$  is the mixing parameter in matter.

Finally, comparing Eqs. 3 and 4, we obtain:

$$\cos 2\theta_m = \frac{\cos 2\theta - \frac{B}{\Delta m^2}}{\sqrt{\left(\cos 2\theta - \frac{B}{\Delta m^2}\right)^2 + \sin^2 2\theta}}$$

$$\sin 2\theta_m = \frac{\sin 2\theta}{\sqrt{\left(\cos 2\theta - \frac{B}{\Delta m^2}\right)^2 + \sin^2 2\theta}}.$$

From which we obtain:

$$\tan 2\theta_m = \frac{\sin 2\theta_m}{\cos 2\theta_m} = \frac{\sin 2\theta}{\cos 2\theta - \frac{B}{\Delta m^2}}.$$