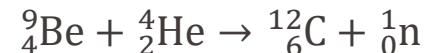


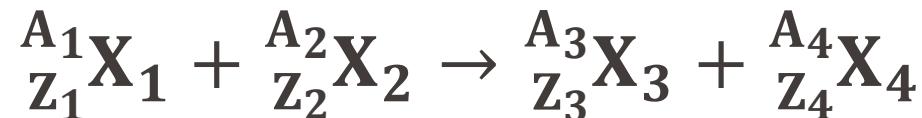
**Nuclear Physics II –
Nuclear Reactions,
Cross-section, Neutron Flux**

- A nuclear reaction occurs when an incoming particle interacts with a nucleus.
- The incoming particle can be **another nuclide** or a **fundamental particle (n, p, α , γ)**.
- **Example** of a common reaction used in neutron sources:

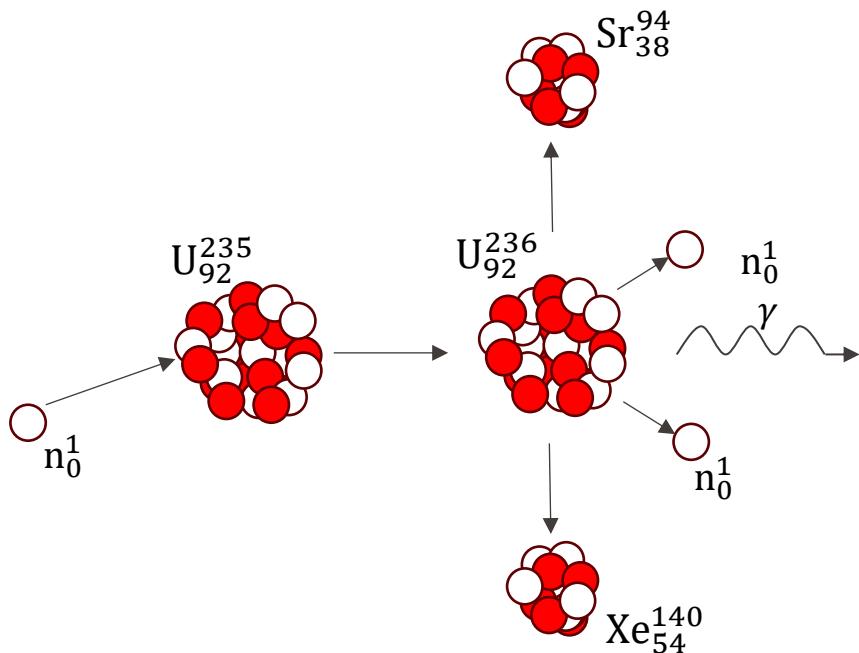
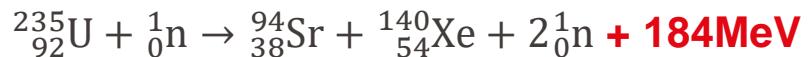


- Similar to chemical reactions but with key differences:
 - Chemical reactions involve the **rearrangement of electrons** in atoms.
 - Nuclear reactions involve **changes within the nucleus**, leading to the **transformation of elements** with **much larger amounts of energy involved** (either released or required).

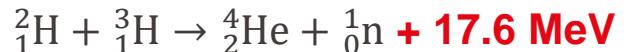
- Nuclear reactions have the general form



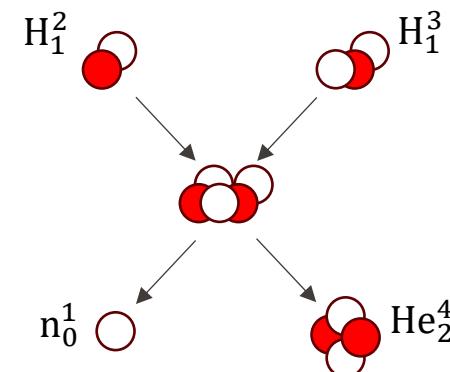
- Main conservation laws to consider:
 - Number of nucleons $\rightarrow A_1 + A_2 = A_3 + A_4$
 - Electric charge $\rightarrow Z_1 + Z_2 = Z_3 + Z_4$
- NOTE:
 - **Alternative notations for reactions involving “basic” particles** $\rightarrow {}_4^9\text{Be}(\alpha, n) {}_{6}^{12}\text{C}$
 - Radioactivity can be considered a particular example of a nuclear reaction with single reactant



- Large **energy density**:
 - ~200 MeV per fission vs. ~4 eV per CH_4 molecule combustion (~50M times more)
- It produces additional neutrons:
 - Each fission neutron can induce further fissions, leading to a self-sustaining **chain reaction**.



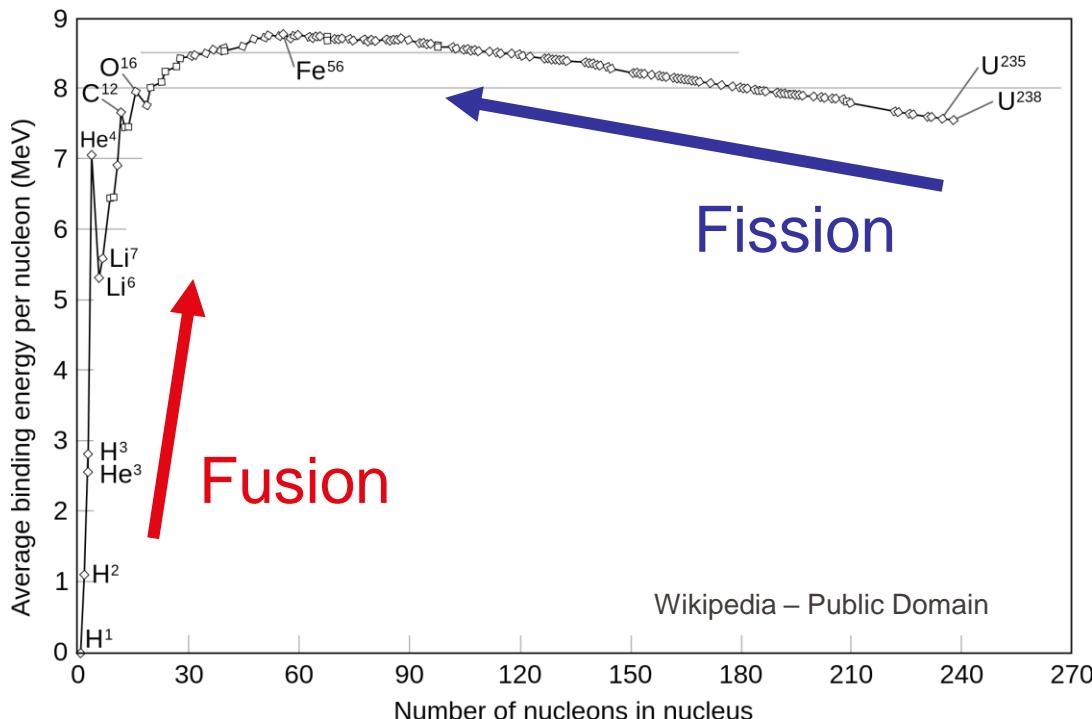
- Even higher energy efficiency than fission:
 - ~200 MeV per fission reaction, but **~0.85 MeV per nucleon**.
 - ~17.6 MeV per fusion reaction, but **~3.5 MeV per nucleon**.
- Need to overcome the Coulomb barrier:
 - In the Sun, **gravitational confinement** creates high pressure enabling fusion at "only" ~15 million K
 - Future reactors will require ~100 million K and **magnetic confinement**.



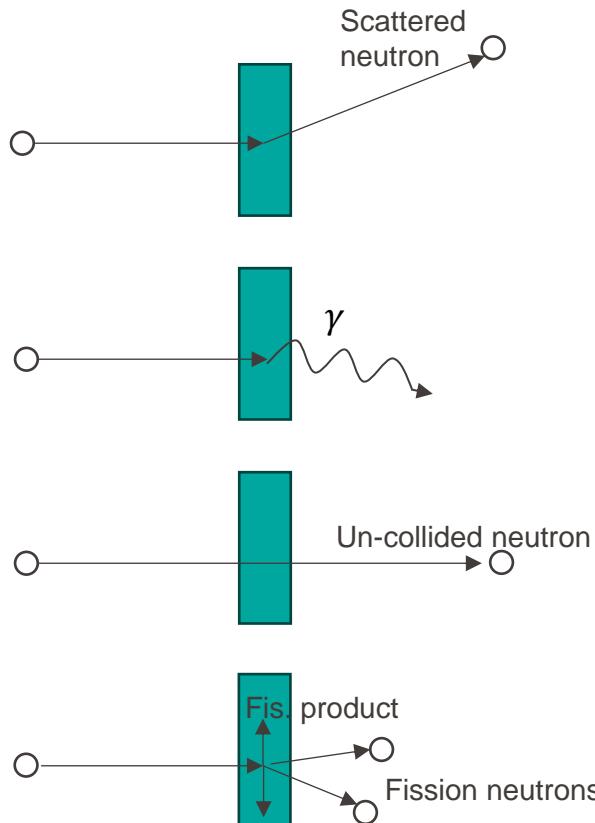
14.1 MeV

3.5 MeV

Average Binding Energy per Nucleon



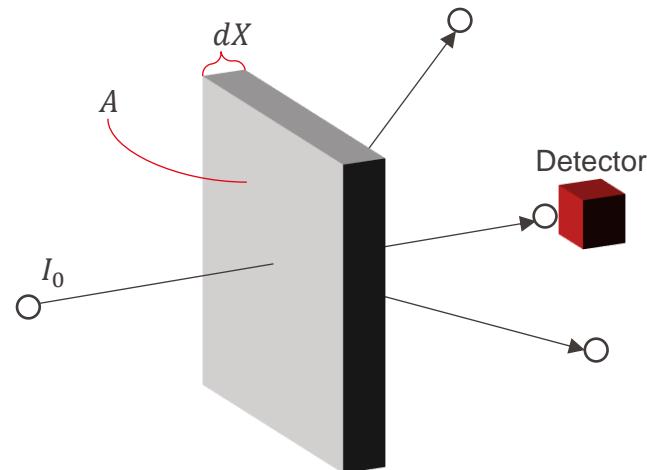
- The mass of a nucleus is slightly less than the sum of its protons and neutrons.
- This difference Δm is the **mass defect** and the **binding energy** is given by:
$$\Delta E = \Delta m \cdot c^2$$
- **$\Delta E/A$ – Binding Energy per Nucleon** measures how strongly nucleons are bound together.
- Higher $\Delta E/A \rightarrow$ More stable nucleus (more energy required to break it apart).
- Reactions resulting in a shift towards the maximum $\Delta E/A$ release energy, i.e. the products are more stable.



- Nuclear reactions are **probabilistic** in nature!
- E.g., a neutron approaching a nucleus can experience different fate:
 - **No interaction!**
 - **Elastic scattering** – The neutron bounces off, losing some energy.
 - **Inelastic scattering** – The neutron excites the nucleus, which emits gamma radiation.
 - **Capture (n,γ)** – The neutron is absorbed, and a gamma ray is emitted, i.e. mass number increases by 1
$${}^A_Z X + {}^1_0 n \rightarrow {}^{A+1}{}_Z X + \gamma$$
- **Fission (n,f)** – The neutron splits the nucleus, releasing energy and more neutrons.
- **Other reactions** – The neutron may induce an (n,α) , (n,p) , $(n,2n)$ or other reactions.

Imagine the following setup:

- A beam of monoenergetic neutrons with density n , speed v and intensity $I_0 = nv$ strikes a thin slab of material.
- The slab has area A , thickness X and nuclide density N .
- A detector measures # of neutrons that pass through $I(X)$



How many reactions or collisions have occurred?

- **Number of reactions:** $-dI \cdot A = (I_0 - I(dX)) \cdot A = \sigma I N A dX$
- **Reaction rate** (per unit volume per second): $R = -\frac{dI}{dx} = \sigma NI = \Sigma I$

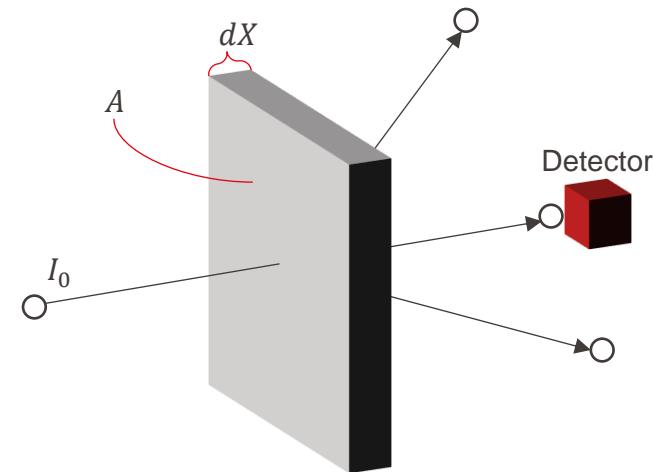
- **Cross-section (sometimes written as XS) quantifies the probability of a neutron interaction.**

Microscopic cross-section σ

- Probability of interaction per nucleus
- Effective area for interaction
- Measured in barns (1 barn = $10^{-24} cm^2$)

Macroscopic cross-section $\Sigma = \sigma N$ (cm^{-1})

- Probability of interaction per cm of material.
- Accounts for all nuclei in the material.
- For a mixture: $\Sigma = \sum_j (N_j \sigma_j)$.



▪ Each reaction type has a different cross-section!

- Total σ_t : sum of scattering and absorption $\sigma_t = \sigma_s + \sigma_a = \underbrace{\sigma_{s,e} + \sigma_{s,i}}_{\text{scattering}} + \underbrace{\sigma_c + \sigma_f + \sigma_p + \sigma_\alpha + \dots}_{\text{absorption}}$

From the definition of reaction rate: $-\frac{dI}{dx} = \sigma NI \rightarrow -\frac{dI}{I} = \sigma N dx$

We can integrate between 0 and x, obtaining the **beam intensity as function of penetration depth**:

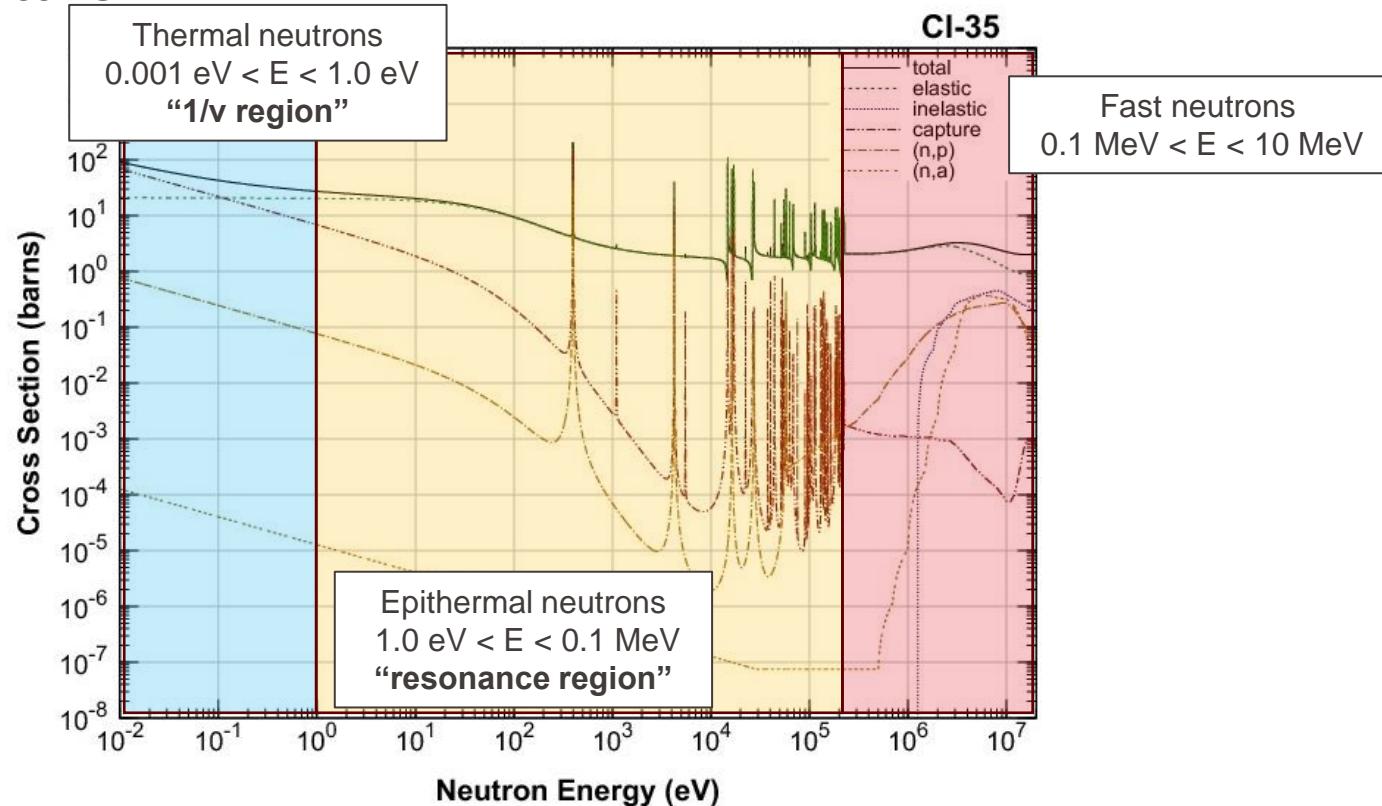
$$I(x) = I_0 \exp(-\Sigma x) = I_0 \exp(-N\sigma x)$$

Probabilistic interpretation:

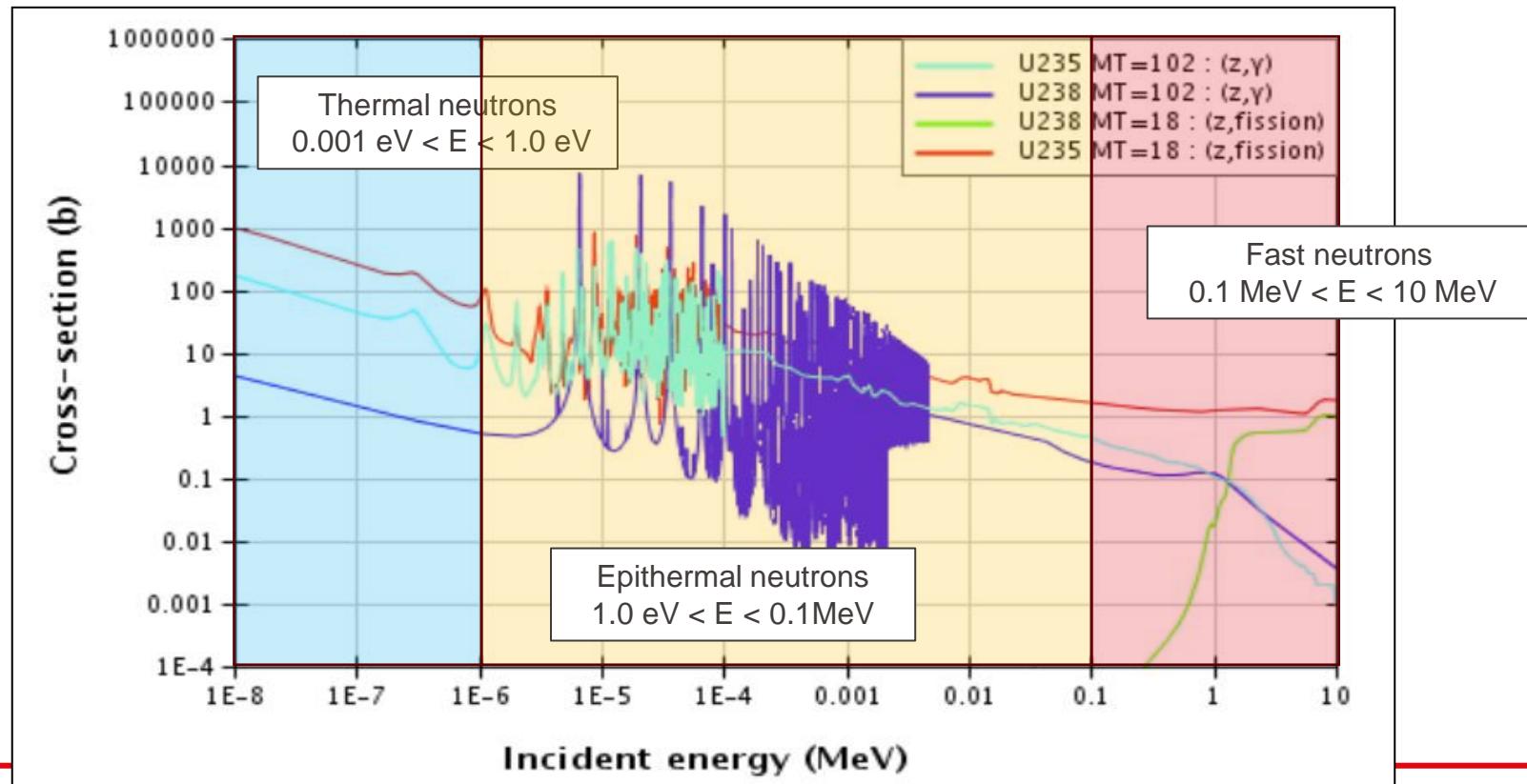
1. $-\frac{dI}{I} = \Sigma dx$ is the probability a neutrons that has survived until x will collide in dx
2. $\frac{I(x)}{I_0} = \exp(-\Sigma x)$ is the probability that a neutron has survived without collision until x
3. From the previous two we can derive that the **mean free path (i.e. the mean distance traveled by a neutron between collisions) is the inverse of the macroscopic cross-section**:

$$\lambda = \int_0^\infty x p(x) = \int_0^\infty x \Sigma \exp(-\Sigma x) dx = 1/\Sigma$$

- Example, Cl-35 XS



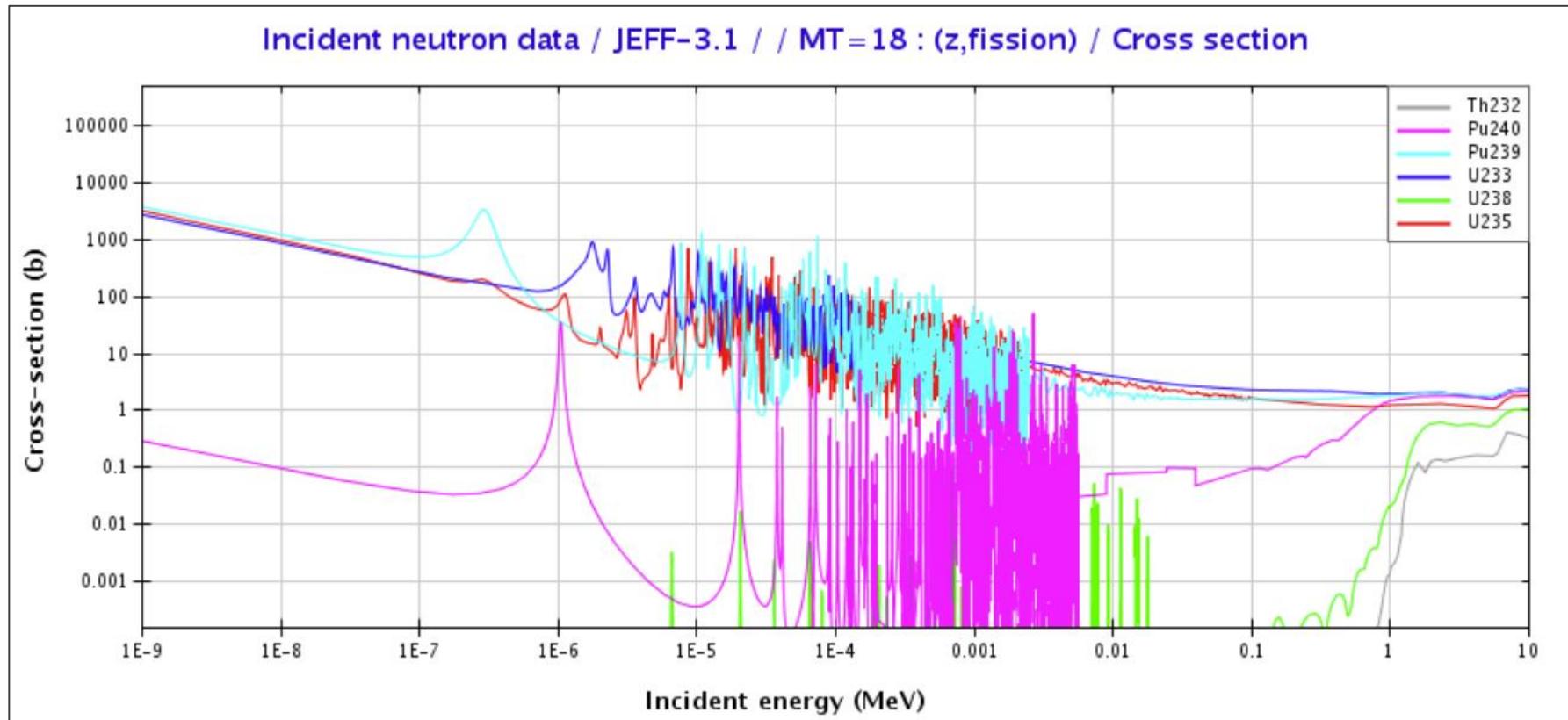
- Example, U-238 and U-238 fission and capture cross-sections

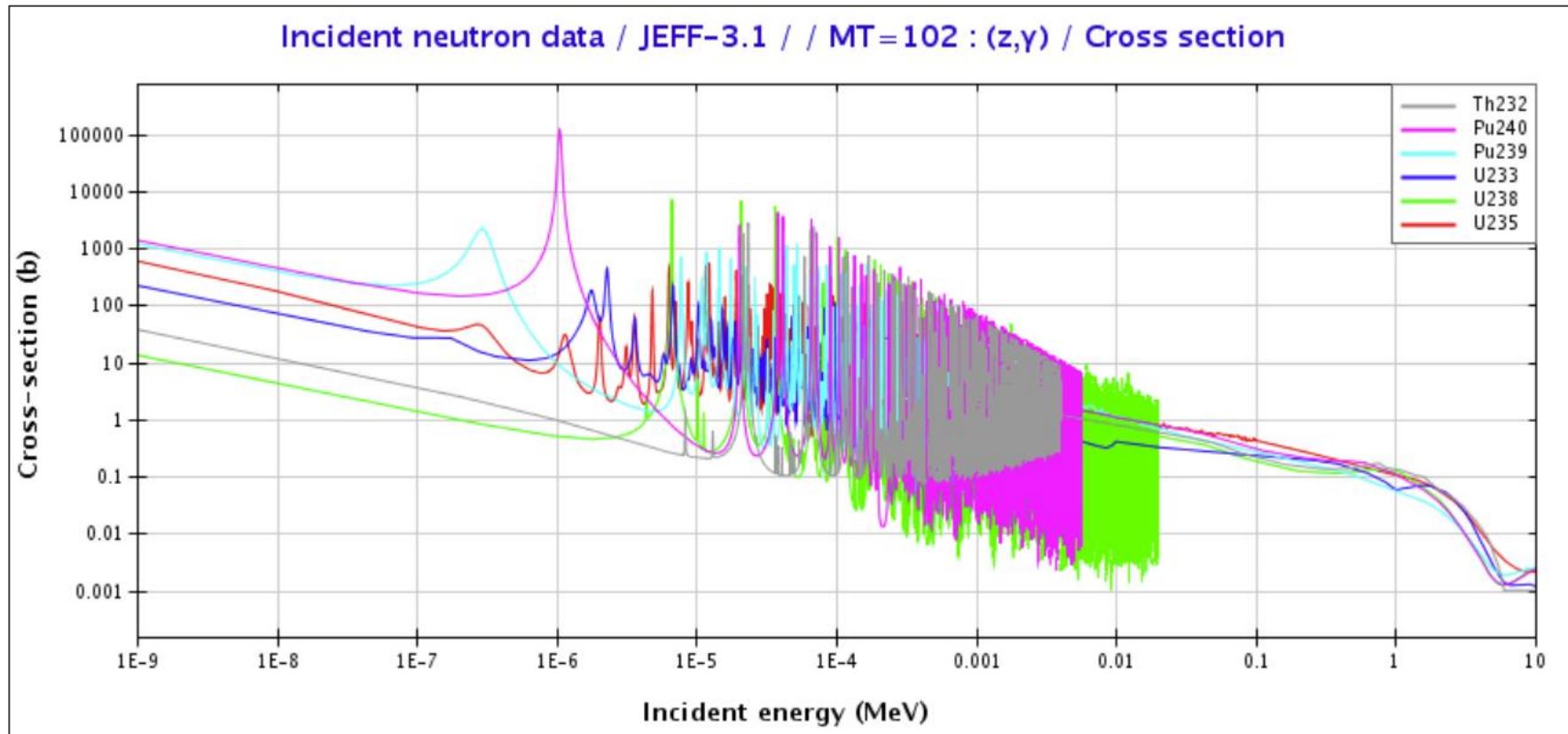


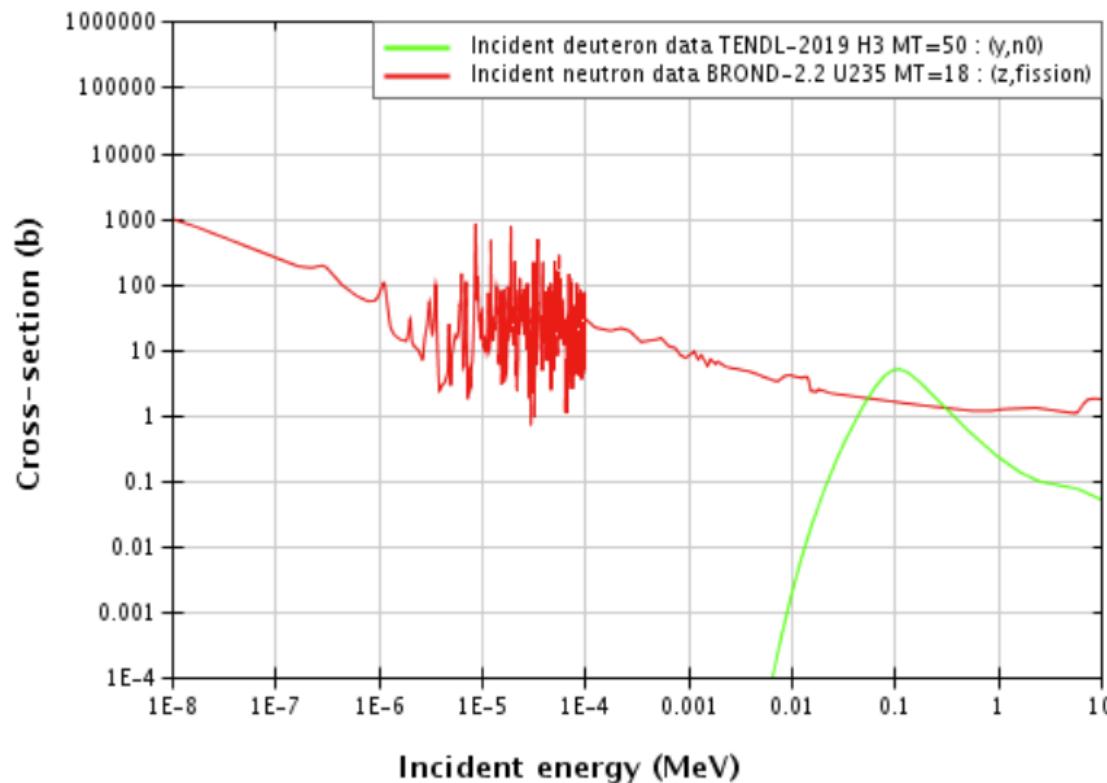
- Nuclear fuel is any **fissionable** material, i.e. a material capable of producing heat from nuclear reactions.
- A fissionable nuclide that can have fission with thermal neutrons with high probability is called **fissile** e.g. ^{233}U , ^{235}U , ^{239}Pu , or ^{241}Pu
- Certain isotopes (often fissionable at high neutron energies) are **fertile**: they provide artificial fissionable isotopes through neutron capture, e.g. ^{232}Th , ^{238}U , or ^{240}Pu



	Fissiles U^{235} , Pu^{239} , U^{233}	Fertiles U^{238} , Th^{232}
Thermal fissions	significant	insignificant
Fast fissions	weak	weak
Captures	parasitic	useful (new fissile is produced)







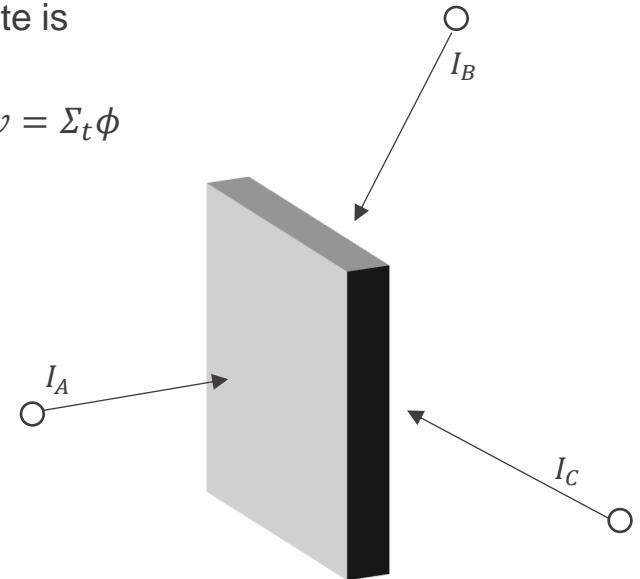
- For neutrons with multiple direction I_A , I_B , etc. the total reaction rate is

$$R = IN\sigma_t = \sum_t (n_A + n_B + \dots) v = \sum_t \cdot \sum n_i v = \sum_t \phi$$

where $\phi = \sum n_i v = nv$ is the **neutron flux and has units of $\frac{\#}{cm^2 s}$**

- The neutron flux can be interpreted as the number of neutrons crossing unit area of the medium in unit time

- Despite the name, the **flux is a scalar (non-directional) quantity.**



- The microscopic XS of a nucleus for a given reaction depends on:
 1. Nothing, it is constant
 2. The material density
 3. The energy of the incoming neutron
 4. Both energy and density

- The microscopic XS of a nucleus for a given reaction depends on:
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 4. Both energy and density
- It is the **macroscopic XS** that depends on medium density!

- A material has a neutron XS of $3.50 \times 10^{-24} \text{ cm}^2/\text{nuclei}$ and contains $4.20 \times 10^{23} \text{ nuclei/cm}^3$
 1. What is the macroscopic cross section?
 2. What is the mean free path?
 3. If neutrons impinge perpendicularly on a slab of the material that is 3.0 cm thick what fraction of them will penetrate the slab without a collision?
 4. What part will collide before 1.5 cm?

$$1. \quad \Sigma = \sigma N = 1.47 \text{ cm}^{-1}$$

$$2. \quad \lambda = \frac{1}{\Sigma} = 0.680 \text{ cm}$$

$$3. \quad \frac{I}{I_0} = e^{-\Sigma x} = 1.21 \%$$

$$4. \quad P = 1 - e^{-\Sigma x} = 89 \%$$

- What is the total macroscopic XS of UO_2 that has been enriched to 4%?
Assume:

- $\sigma_{25} = 607.5 \text{ b}$
- $\sigma_{28} = 11.8 \text{ b}$
- $\sigma_O = 3.8 \text{ b}$
- $\rho_{\text{UO}_2} = 10.5 \text{ g/cm}^3$

1. Compute the total microscopic XS of uranium (4%)

$$\sigma_U = \epsilon * 607.5 + (1 - \epsilon) * 11.8 = 35.63 \text{ b}$$

2. Compute the microscopic XS of UO_2 (two nuclei of oxygen per uranium)

$$\sigma_{\text{UO}_2} = \sigma_U + 2 * \sigma_O = 43.23 \text{ b}$$

3. Compute the macroscopic XS of UO_2

$$\Sigma_{\text{UO}_2} = \frac{\rho_{\text{UO}_2} * N_A}{238 + 2 * 16} * \sigma_{\text{UO}_2} = 1.06 \text{ cm}^{-1}$$

- A nuclear reactor of cylindrical shape ($H = 3.5 \text{ m}$, $R = 1.5 \text{ m}$) operates at 1 GW. Compute the macroscopic XS of fission if the neutron flux is homogeneous and equal to $10^{13} \text{ n/cm}^2/\text{s}$

- $\phi = 10^{13} \text{ cm}^{-2}\text{s}^{-1}$
- $E_f = 200 \text{ MeV}$ ($1 \text{ eV} = 1.6 * 10^{-19} \text{ J}$)
- $P = 1 \text{ GW}$

1. Compute the reaction rate

$$R.R. = \frac{P}{E_f V} = 1.26 * 10^{12} \frac{\text{fissions}}{\text{cm}^3}$$

2. Compute the macroscopic XS

$$\Sigma_f = \frac{R.R.}{\phi} = 0.126 \text{ cm}^{-1}$$

- A reactor with same dimension and power of that of Ex.4, operates instead with a non-homogeneous flux, given by

$$\phi(x, R) = \phi_0 \cos\left(\frac{\pi x}{H}\right) J_0\left(\frac{a_{01}r}{R}\right)$$

where x varies between $-H/2$ and $H/2$ and

- Compute the new fission total XS if

- $\phi_0 = 1 * 10^{19} \text{ cm}^{-2} \text{ s}^{-1}$
- $\int_0^R J_0\left(\frac{a_{01}r}{R}\right) dr = \frac{R^2}{a_{01}^2}$, $a_0 = 2.405$

- Compute also the uranium enrichment assuming that the reactor core's volume is 55.4% fuel and the rest is water and if
 - $\sigma_{f,U} = 580 \text{ b}$

1. Compute the total R.R

$$R.R. = \frac{P}{E_f} = 3.125 * 10^{19} \frac{\text{fission}}{\text{s}}$$

2. Compute the total macroscopic XS

$$R.R. = \int_V \Sigma_f \phi dV = \Sigma_f \int_V \phi dV$$

$$\int_V \phi dV = \phi_0 \int_{2\pi} d\theta \int_0^R J_0 \left(\frac{a_{01} r}{R} \right) dR \int_{-\frac{H}{2}}^{\frac{H}{2}} \cos \left(\frac{\pi x}{H} \right) dx$$

$$\Sigma_f = \frac{R.R.}{\int_V \phi dV} = 0.57 \text{ cm}^{-1}$$

2. Compute the uranium enrichment

$$\Sigma_{f,tot} = 0.554 * \Sigma_{f,UO_2} + (1 - 0.554) * \Sigma_{f,H_2O} = 0.544 * \Sigma_{f,UO_2}$$

$$\Sigma_{f,UO_2} = \rho_{UO_2} * \frac{N_A}{238 + 2 * 16} * (\epsilon \sigma_{f,235} + (1 - \epsilon) \sigma_{f,238} + 2 * \sigma_{f,O})$$

$$\epsilon = \frac{\Sigma_{f,tot} * 0.554}{\rho_{UO_2} * \frac{N_A}{238 + 2 * 16} * \sigma_{f,235}} = 2.32\%$$