

Black body radiation

Definition

A system that completely absorbs all incident electromagnetic energy. This absorption manifests itself as thermal agitation, which in turn gives rise to thermal radiation.

Rationale


The radiation emitted by a star can be approximated by that of a black body.

A star interior is a very thick medium, a source of energy through thermonuclear fusion.

However, the radiation is effectively “trapped” within the stellar interior by numerous interaction processes.

On average, photons require millions of years to escape. Consequently, the stellar medium behaves analogously to the cavity used to model a black body.

Planck Law

$$I_\nu = B_\nu(T) = \frac{2h\nu^3}{c^2} \frac{1}{e^{\frac{h\nu}{kT}} - 1}.$$


$$I_\nu d\nu = I_\lambda d\lambda$$

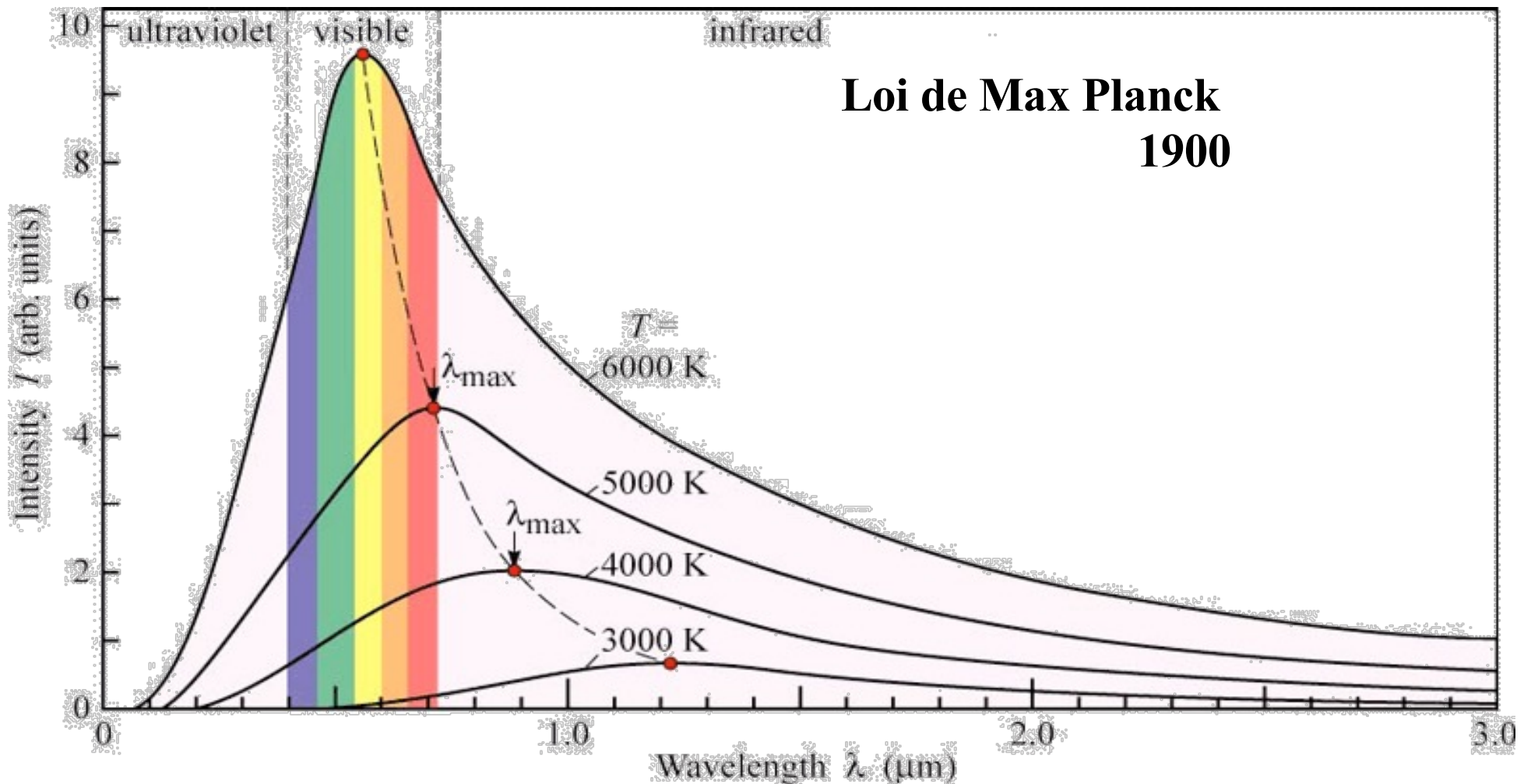
$$I_\lambda / I_\nu = |d\nu / d\lambda|$$

$$I_\lambda = \frac{2hc^2}{\lambda^5} \frac{1}{e^{hc/kT\lambda} - 1}$$

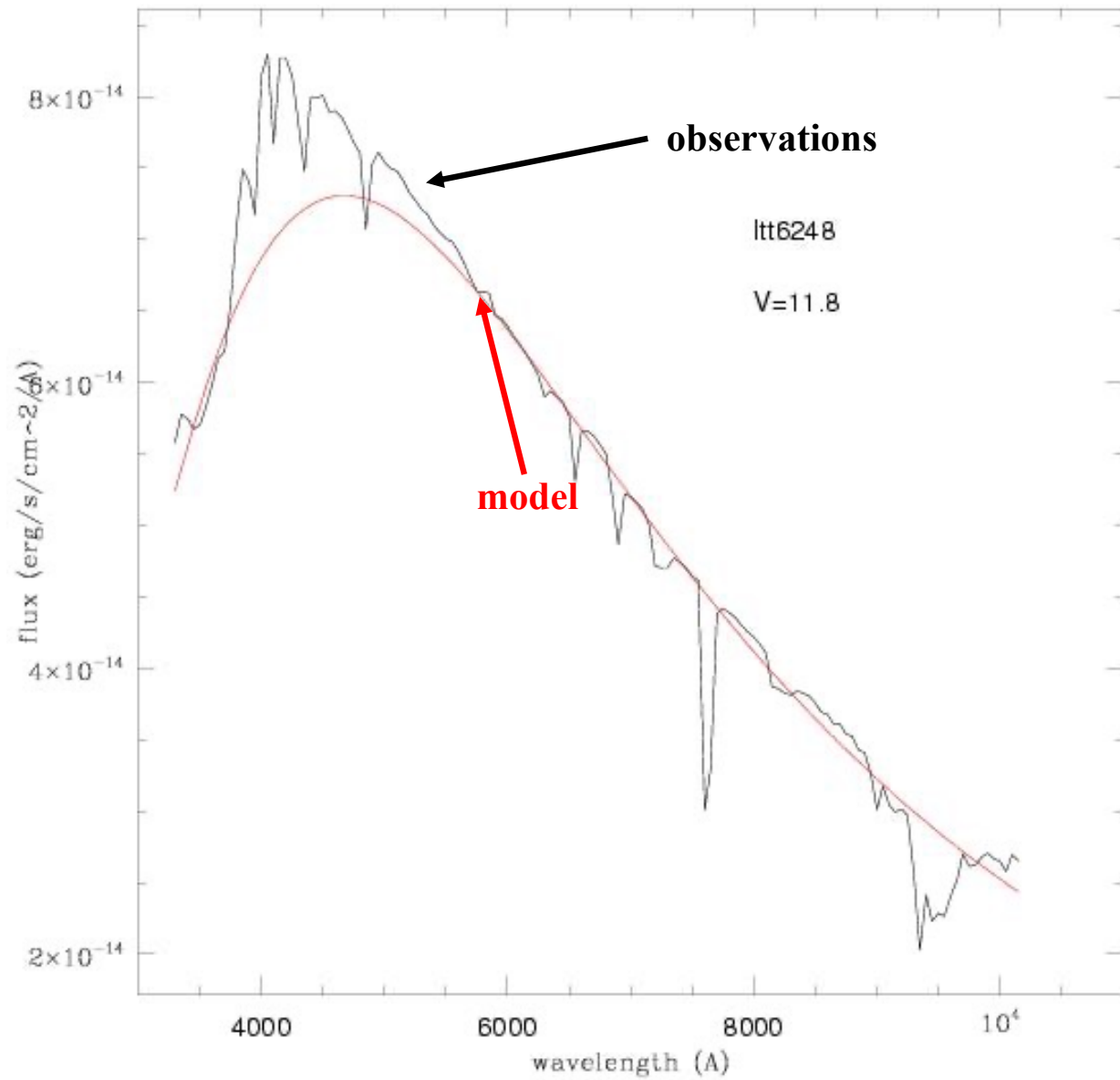
Boltzmann constant $k = 1.3805 \cdot 10^{-23}$ J/K

Planck constant $h = 6.62 \cdot 10^{-34}$ Js

Emission from black bodies at different temperatures

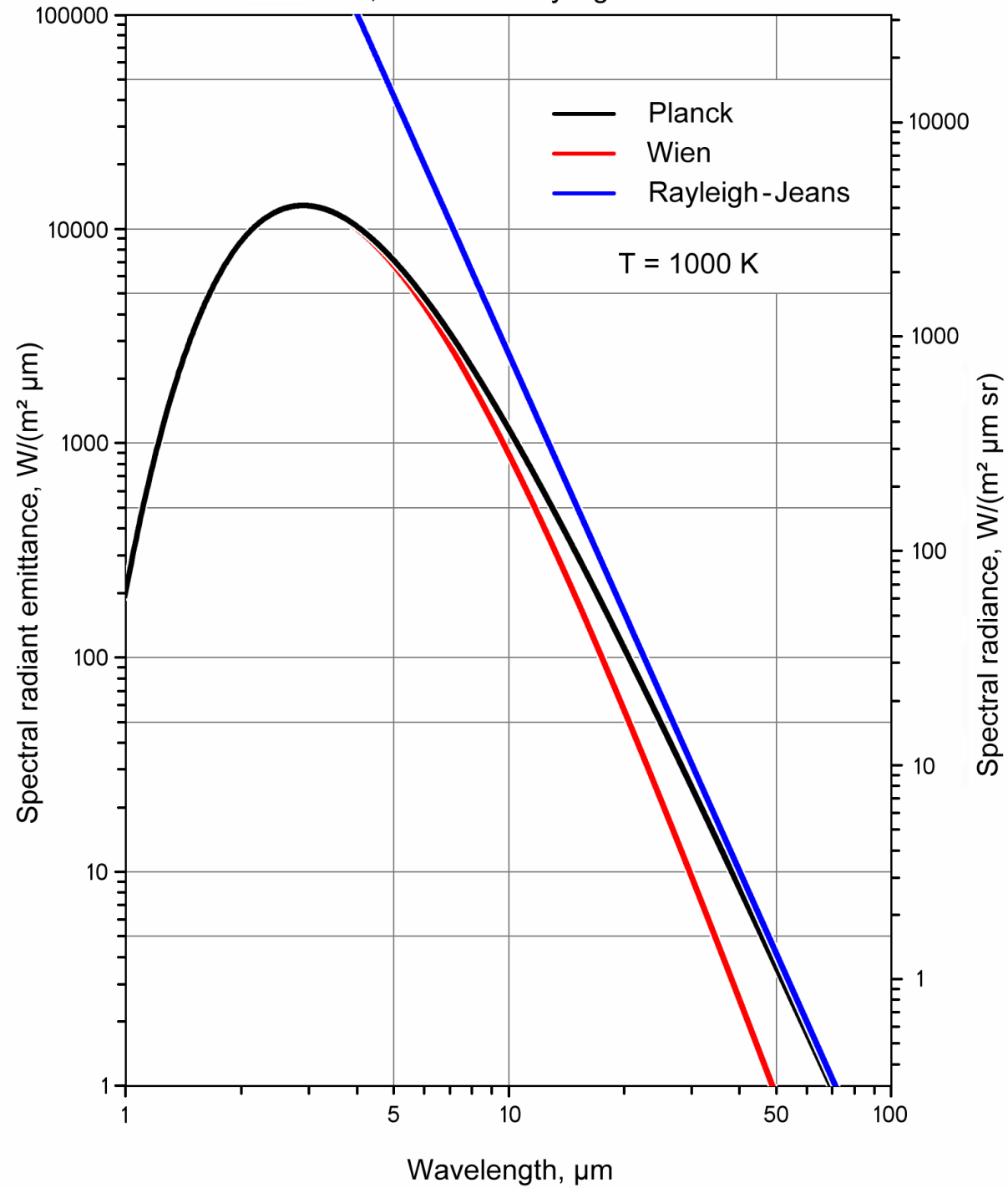


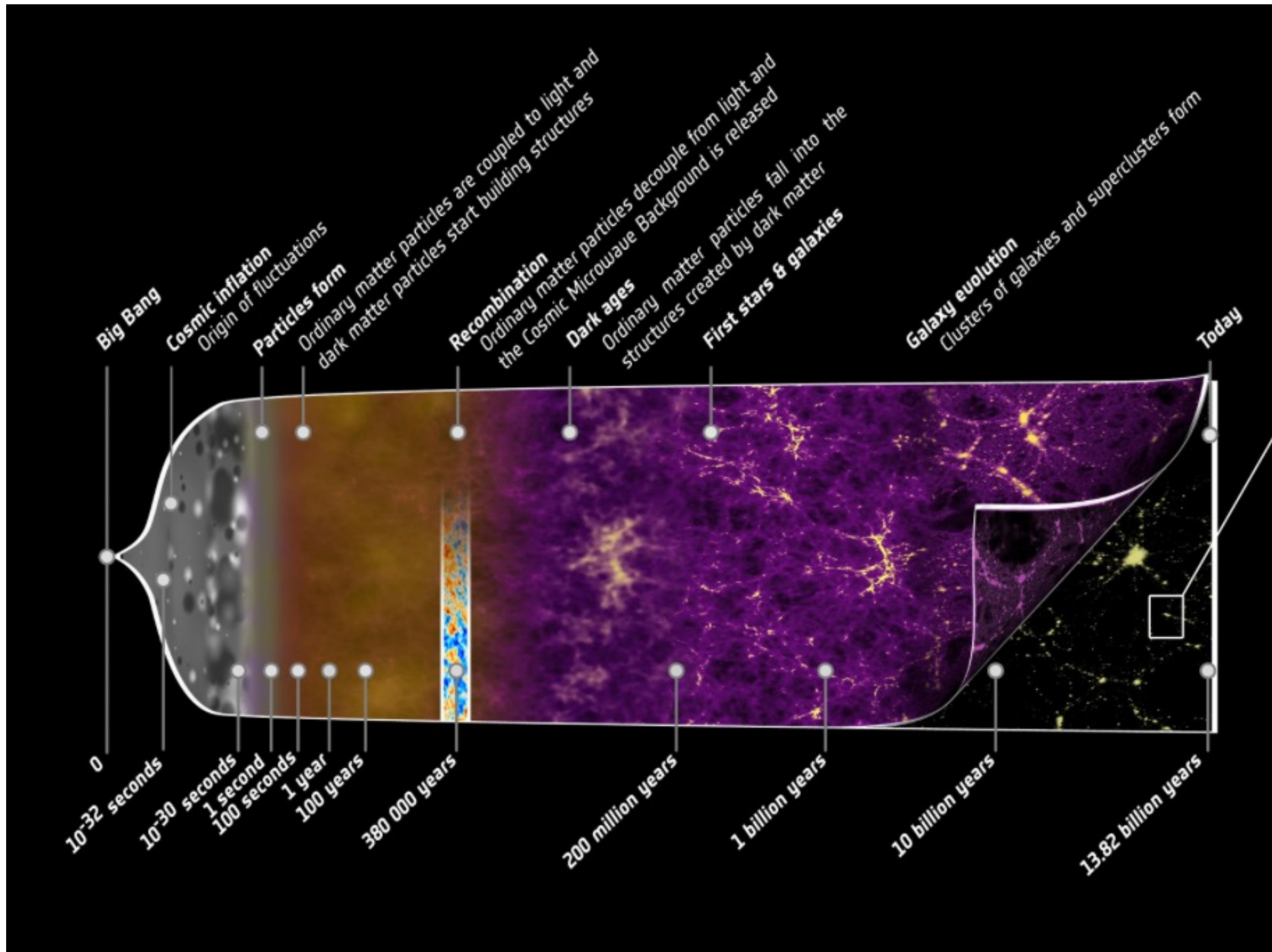
Black body analogous for stars



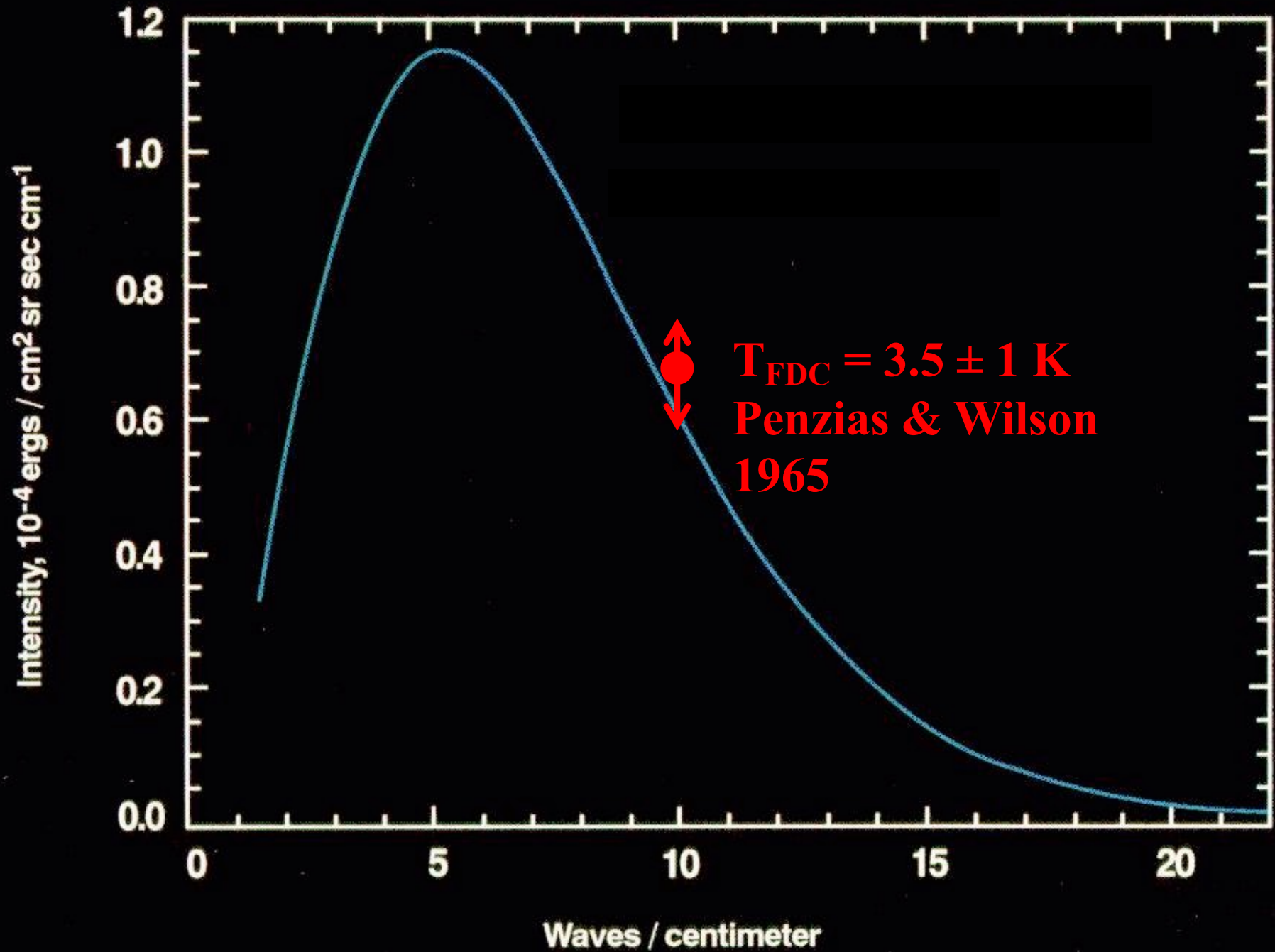
Fit at T = 5350 K

Thermal radiation spectra according to
Planck, Wien and Rayleigh-Jeans





COSMIC MICROWAVE BACKGROUND SPECTRUM



COSMIC MICROWAVE BACKGROUND SPECTRUM FROM COBE

