

Astrophysics II - Exercise Sheet 9

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1 Saha equation, partially ionized medium

Consider a medium composed of a mixture of hydrogen and helium and in which the temperature is $T = 12500$ K and the density $\rho = 10^{-7}$ kg m⁻³. We will neglect radiation pressure. The abundances of hydrogen and helium (by number) are $\nu_1 = 0.9$ and $\nu_2 = 0.1$ respectively. Using the Saha law, we will construct a system of equations that can be solved by iteration to determine the ionization state of the medium. We first define

$$E \equiv \frac{\text{Number of free electrons}}{\text{Total number of atoms or ions}} \quad (1)$$

as well as the quantities

$$X_i^s \equiv \frac{\text{Number of atoms } i \text{ in ionization state } s}{\text{Total number of atoms or ions of species } i} \quad (2)$$

Such that we have

$$E = \sum_i \nu_i \sum_{s=0}^{Z_i} s X_i^s \quad (3)$$

In our particular case,

$$X_1^1 = \frac{\text{H}^+}{\text{H} + \text{H}^+}, \quad X_2^1 = \frac{\text{He}^+}{\text{He} + \text{He}^+ + \text{He}^{++}}, \quad X_2^2 = \frac{\text{He}^{++}}{\text{He} + \text{He}^+ + \text{He}^{++}}, \quad (4)$$

where the fractions are intended by number. For simplicity, we will simply denote these quantities by X_1, X_2, X_3 , in order.

1. Write down the equation giving E in the system considered here.

The Saha equation yields

$$\frac{X_1 P_e}{(1 - X_1) P_g} = K_1, \quad (5)$$

where

$$K_1 = \frac{(2\pi m_e)^{3/2} (kT)^{5/2}}{\beta P h^3} \omega_1 e^{-I_1/kT}. \quad (6)$$

P_e is the electron gas pressure, P_g is the gas pressure (including atoms, ions and electrons) and $\beta \equiv P_g/P$ is the ratio of the gas pressure to the total pressure of the medium. I_1 is the energy difference of the transition $\text{H} \rightarrow \text{H}^+$ and

$$\omega_1 = 2 \frac{U_{\text{H}^+}}{U_{\text{H}}}. \quad (7)$$

Furthermore, we have seen in the lecture that the gas pressure can be related to density through

$$P_g = \frac{k}{m_H} \frac{1 + E}{\mu_0} \rho T, \quad (8)$$

where μ_0 is the mean molecular weight of the atom and ion mixture, while for the electron pressure

$$P_e = \frac{k}{m_H} \frac{E}{\mu_0} \rho T. \quad (9)$$

Combining all of this and adding the Saha equations for the two other ionic species, we get the following system,

$$\left. \begin{aligned} \frac{X_1}{1 - X_1} \frac{E}{1 + E} &= K_1 \\ \frac{X_2}{1 - X_2 - X_3} \frac{E}{1 + E} &= K_2 \\ \frac{X_3}{X_2} \frac{E}{1 + E} &= K_3 \end{aligned} \right\} \quad (10)$$

which is closed by the equation for E you found in (1). The K_i are defined equivalently to K_1 .

2. Solve this system of equations by iteration, by starting from the initial guess that hydrogen is fully ionized and helium completely neutral. *Hint:* use this starting point to obtain an initial guess for E , and move forward from there. You can either do this "by hand", or write a short Python script to automate the process. If you choose to do the latter, try varying the temperature of the medium over the range [5'000, 40'000] K and plot the fractions of He^+ and He^{++} to see how the ionization state depends on temperature.

Indications

$$I_1 = 13.598 \text{ eV}$$

$$I_2 = 24.587 \text{ eV}$$

$$I_3 = 54.416 \text{ eV}$$

$$\omega_1 = 2 \frac{U_H^+}{U_H} = 2 \cdot \frac{1}{2} = 1$$

$$\omega_2 = 2 \frac{U_{He^+}}{U_{He}} = 2 \cdot \frac{2}{1} = 4$$

$$\omega_3 = 2 \frac{U_{He^{++}}}{U_{He^+}} = 2 \cdot \frac{1}{2} = 1$$

$$k = 1.3807 \cdot 10^{-23} \text{ J K}^{-1}$$

$$h = 6.626 \cdot 10^{-34} \text{ J S}$$

$$m_e = 9.109 \cdot 10^{-31} \text{ kg}$$

$$m_H \simeq u = 1.661 \cdot 10^{-27} \text{ kg}$$